

Giant planets near the 5:2 resonance and their influence on terrestrial planets in the habitable zone

E. Pilat-Lohinger (Univ.Vienna)
Aron Süli (Univ.Budapest)
Philippe Robutel (IMCCE)
Florian Freistetter (Univ. Jena)

Motivation

A study of
Extra-solar planetary systems
similar to
our solar system

Classification of the known multi-planet systems (S.Ferraz-Mello, 2005)

- **Class Ia** → Planets in mean motion resonance
- **Class Ib** → Low-eccentricity near-resonant planet pairs
- **Class II** → Non-resonant planets with significant secular dynamics
- **Class III** → Hierarchical planet pairs

Initial Conditions and Computations

Jupiter: on its orbit

Saturn: $a_{sat} = 8 \dots 11$ AU
 $m_{sat} = 1 \dots 30 \times m_{Sat}$

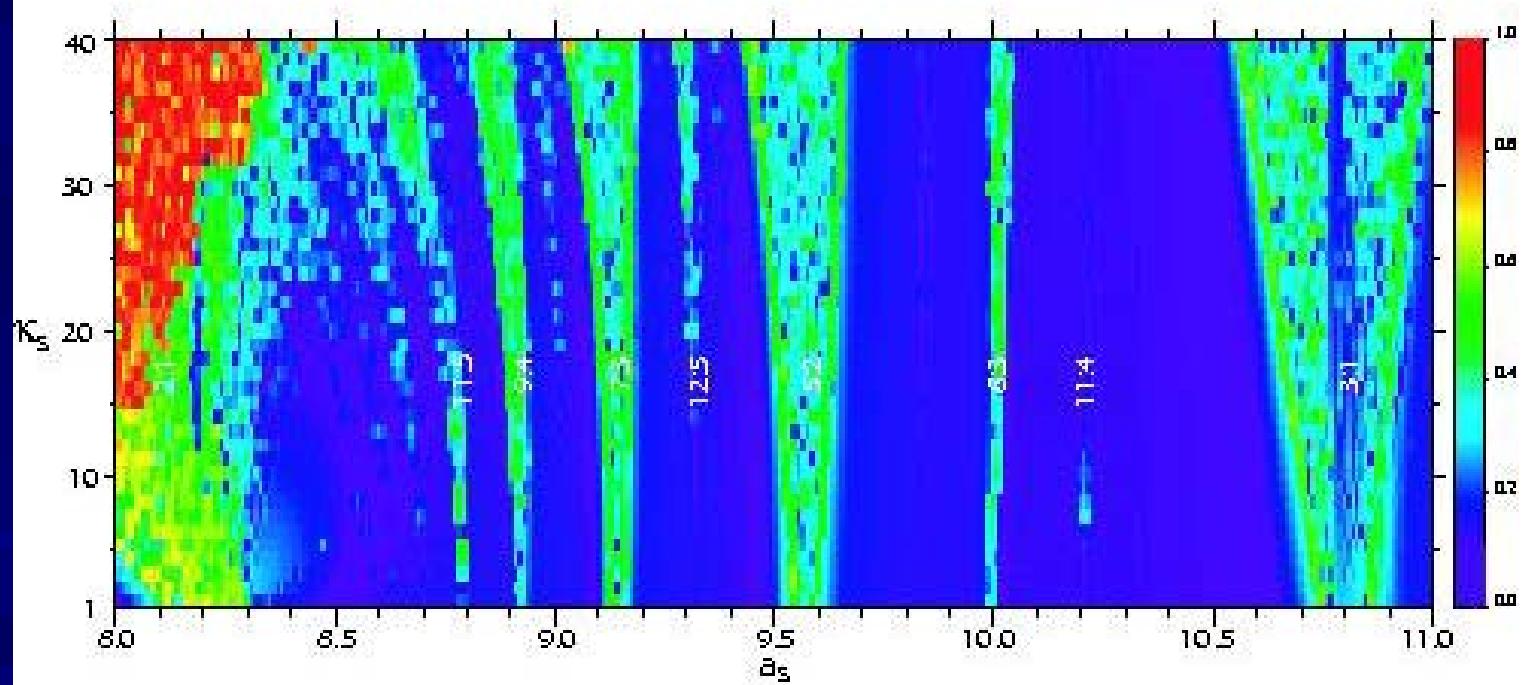
Testplanets in the HZ:

$a_{tp} = 0.6 \dots 1.6$ AU

Mercury 6 (J. Chambers)

Integration time:
20 mio years

HZ: maximum ecc.



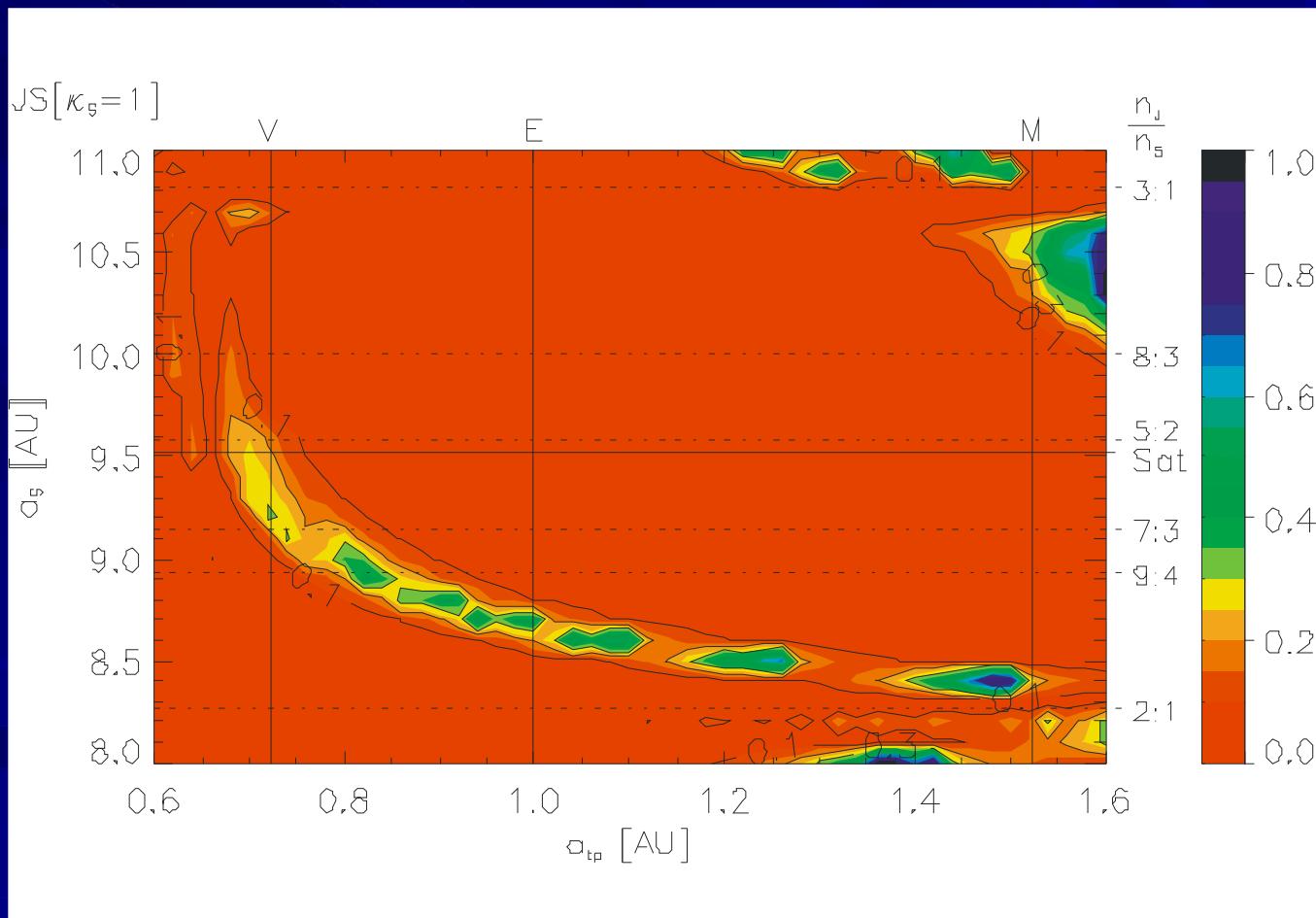
HZ im Sonnensystem:

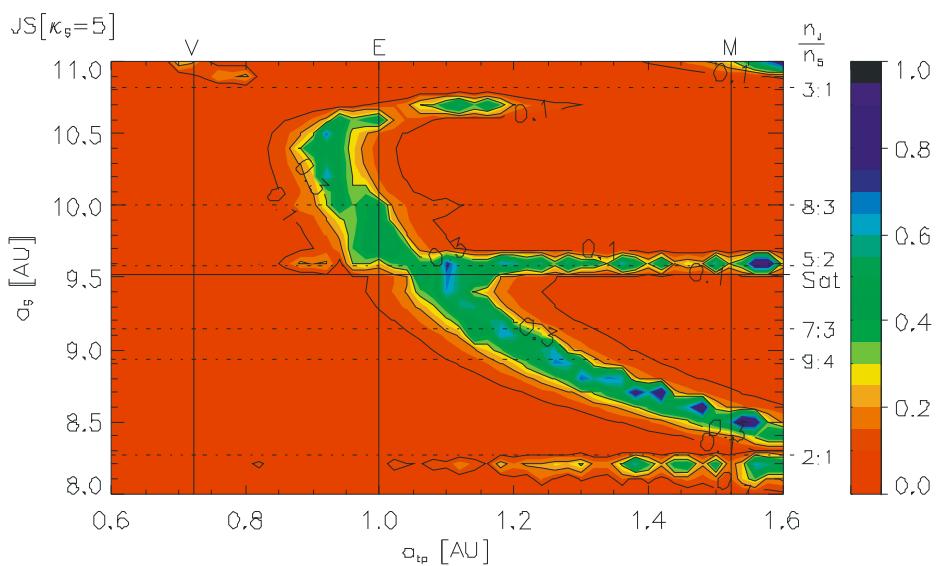
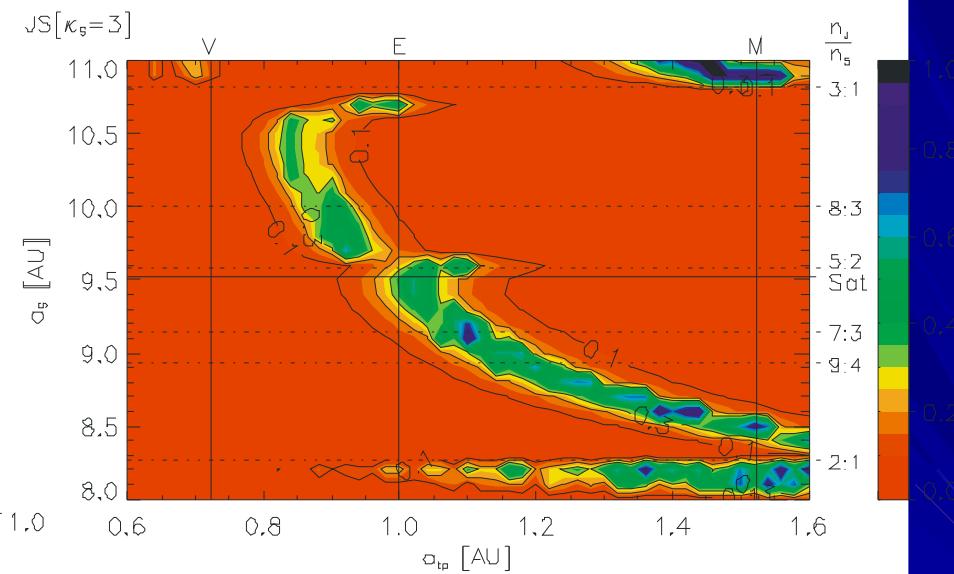
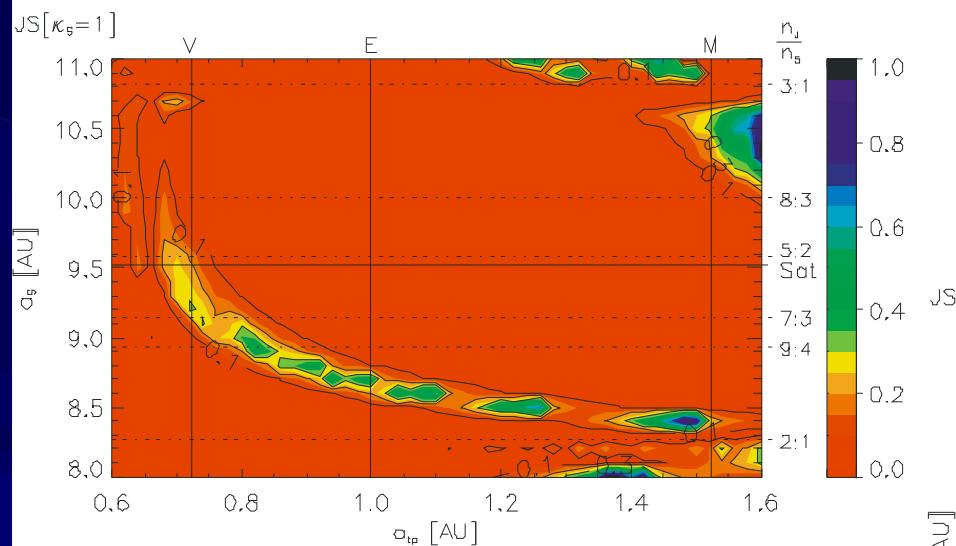
- Kasting: 0.93 – 1.3 AU
- Mischna: 0.93 – 1.7 AU
- Forget: 0.93 – 2 AU

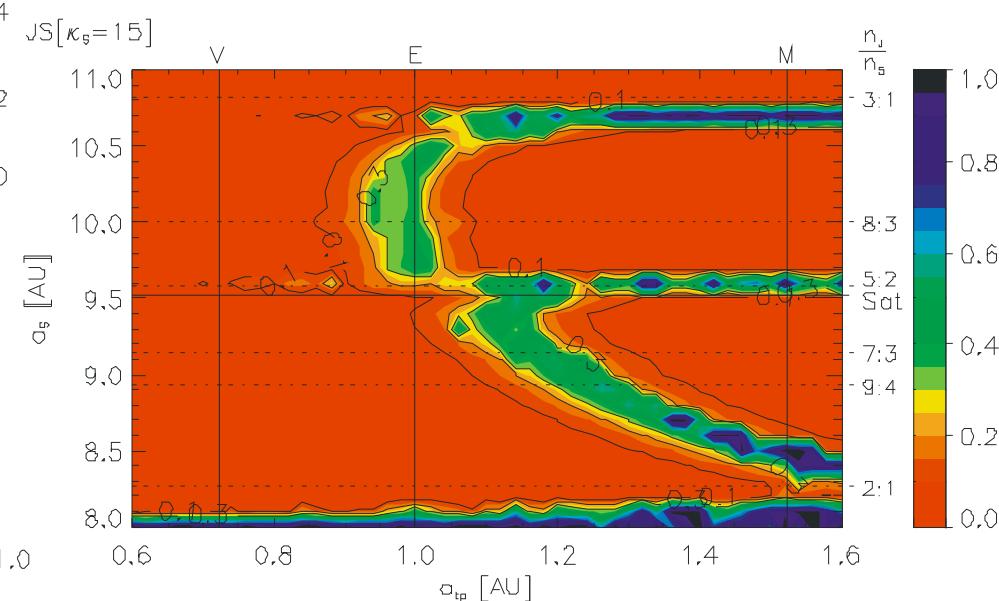
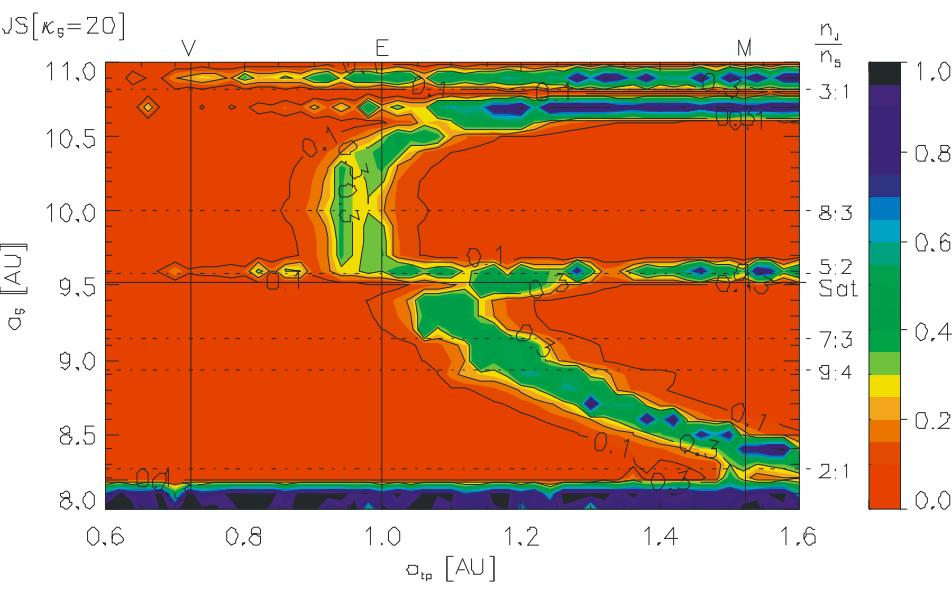
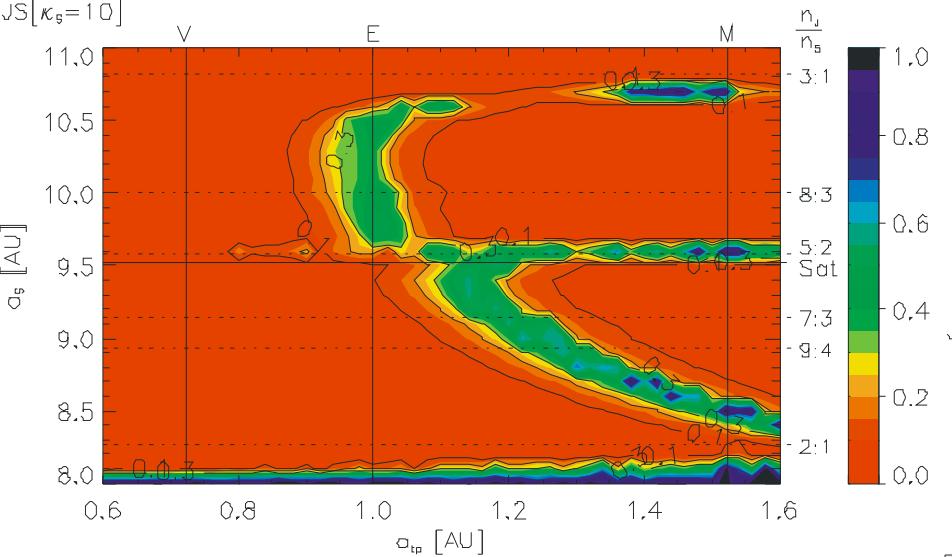
a < 0.93 AU → H₂O becomes a major atmospheric compound and is rapidly lost to space after UV photolysis

a > 1.3 AU → CO₂ condensates in the atmosphere producing CO₂-clouds, that can affect significantly the T-CO₂ coupling

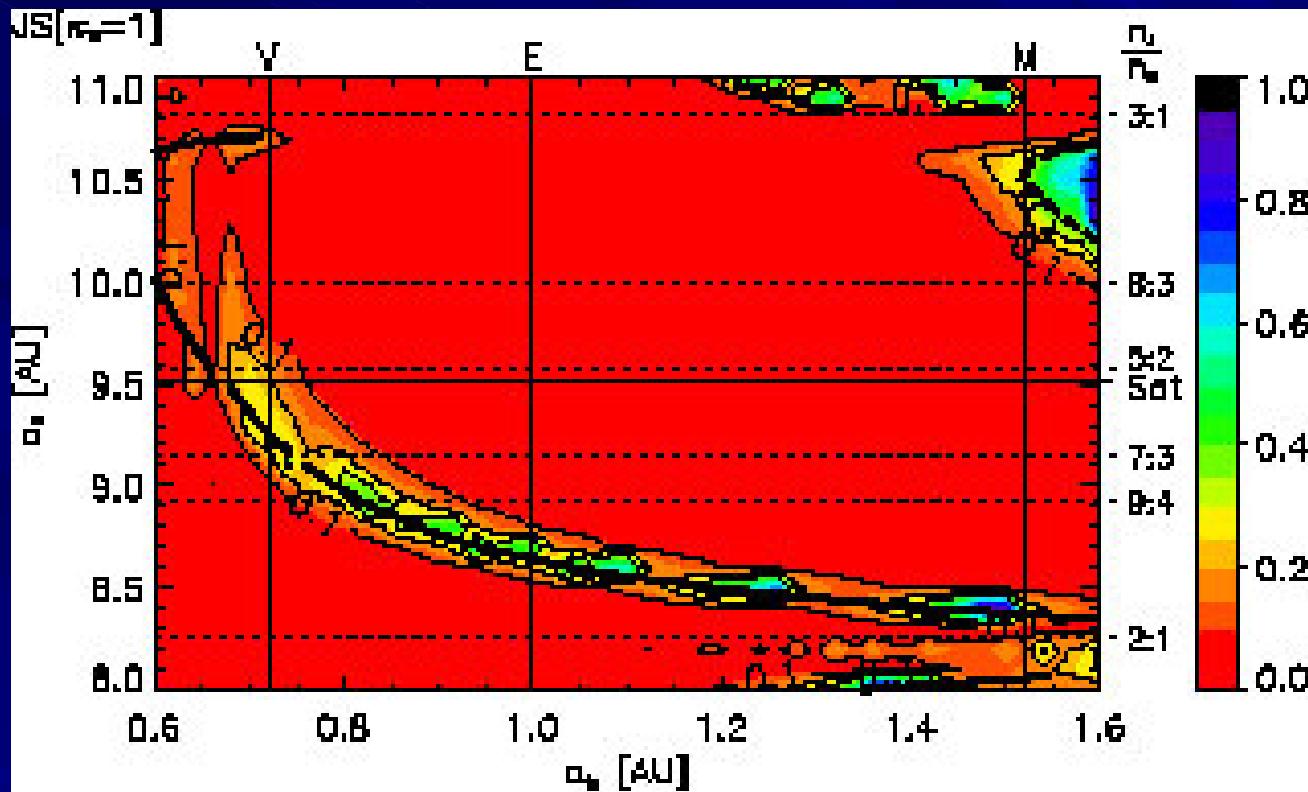
Sun – Jupiter – Saturn



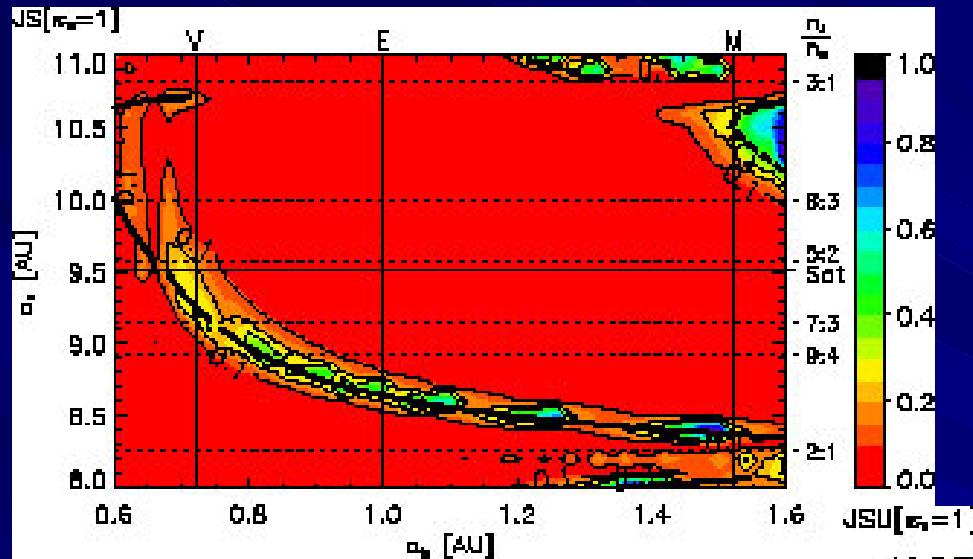




Sun – Jupiter – Saturn

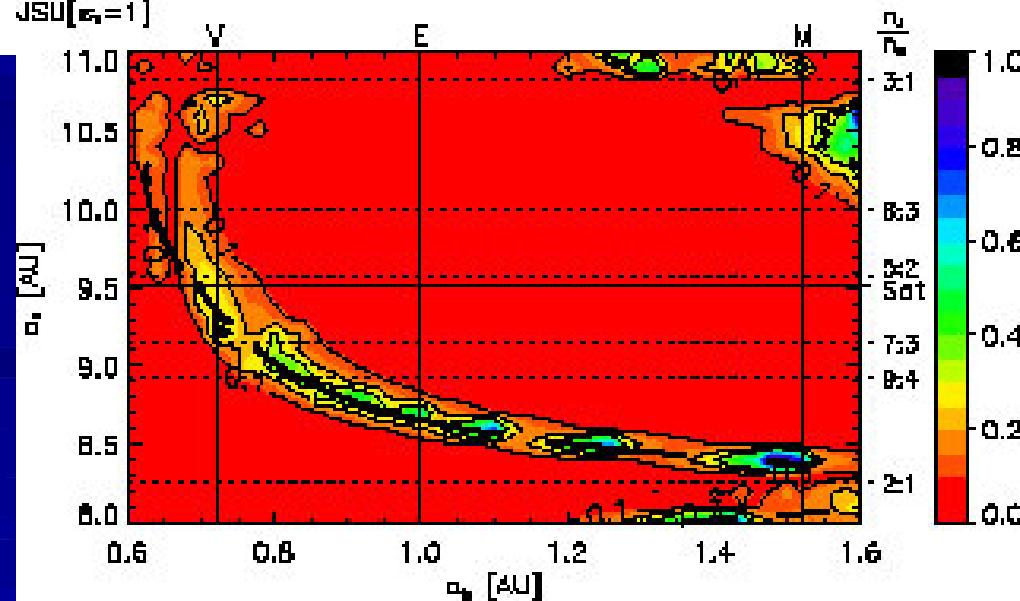


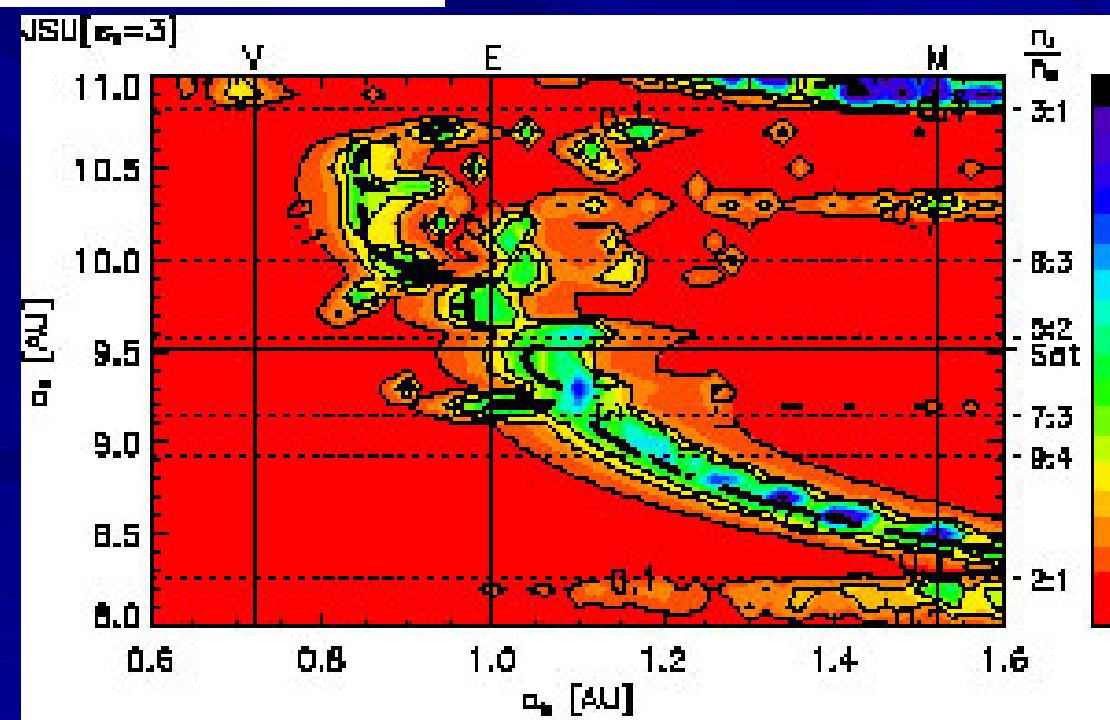
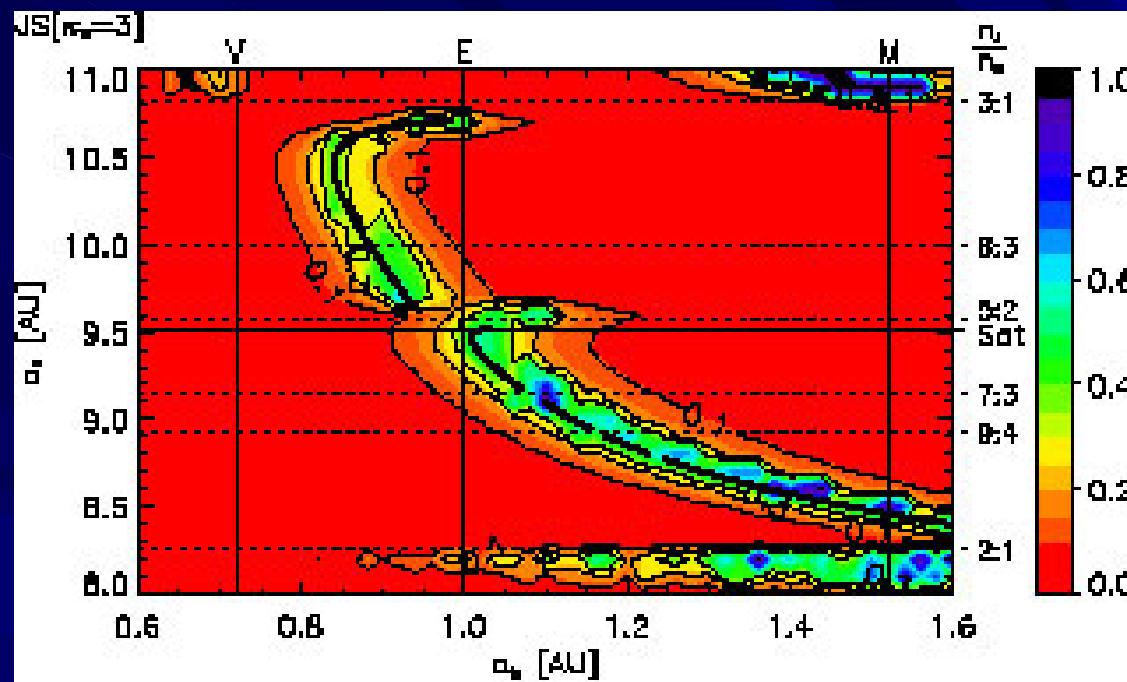
Influence of a third giant planet

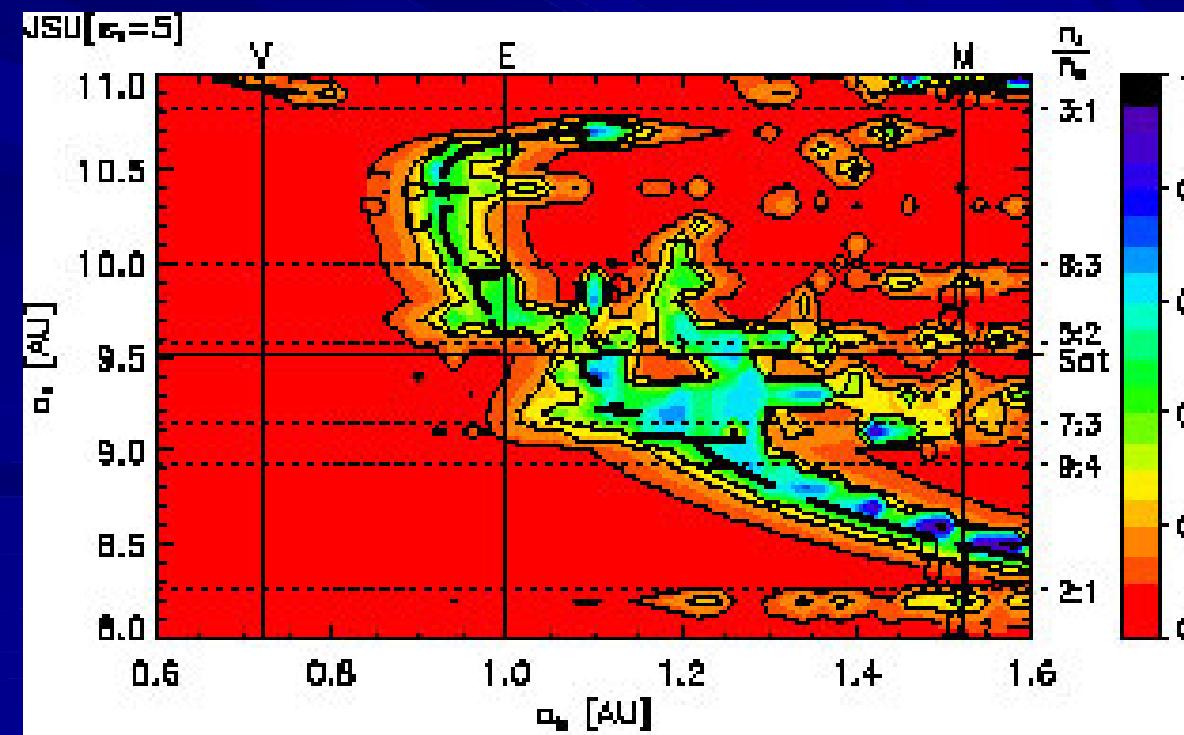
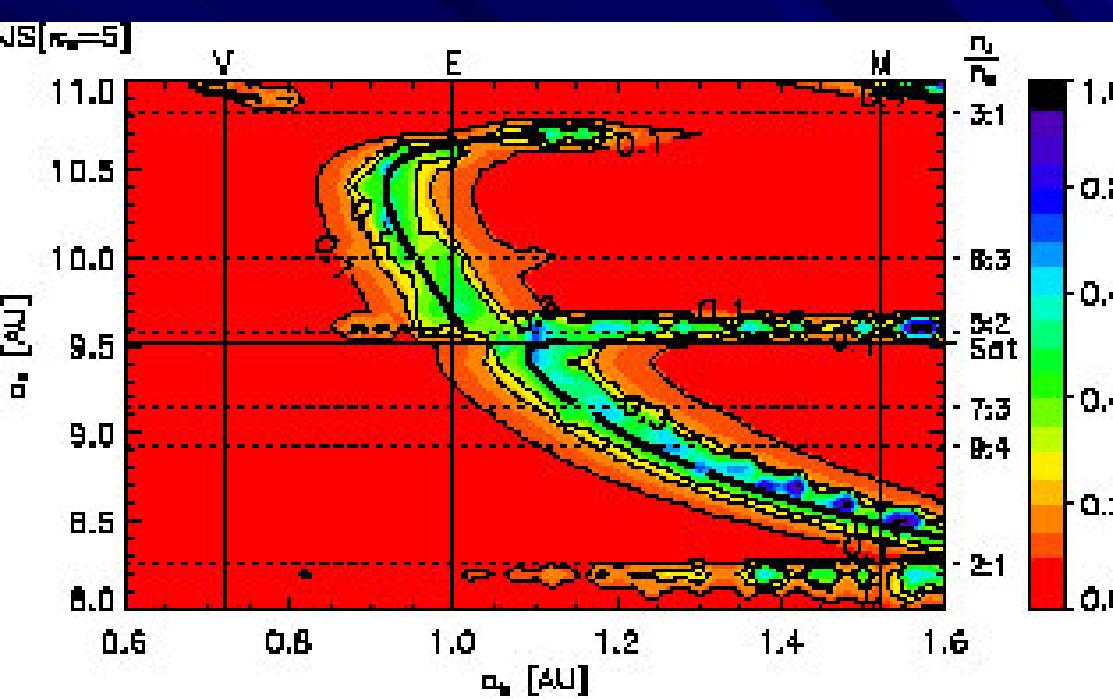


Jupiter – Saturn -Uranus

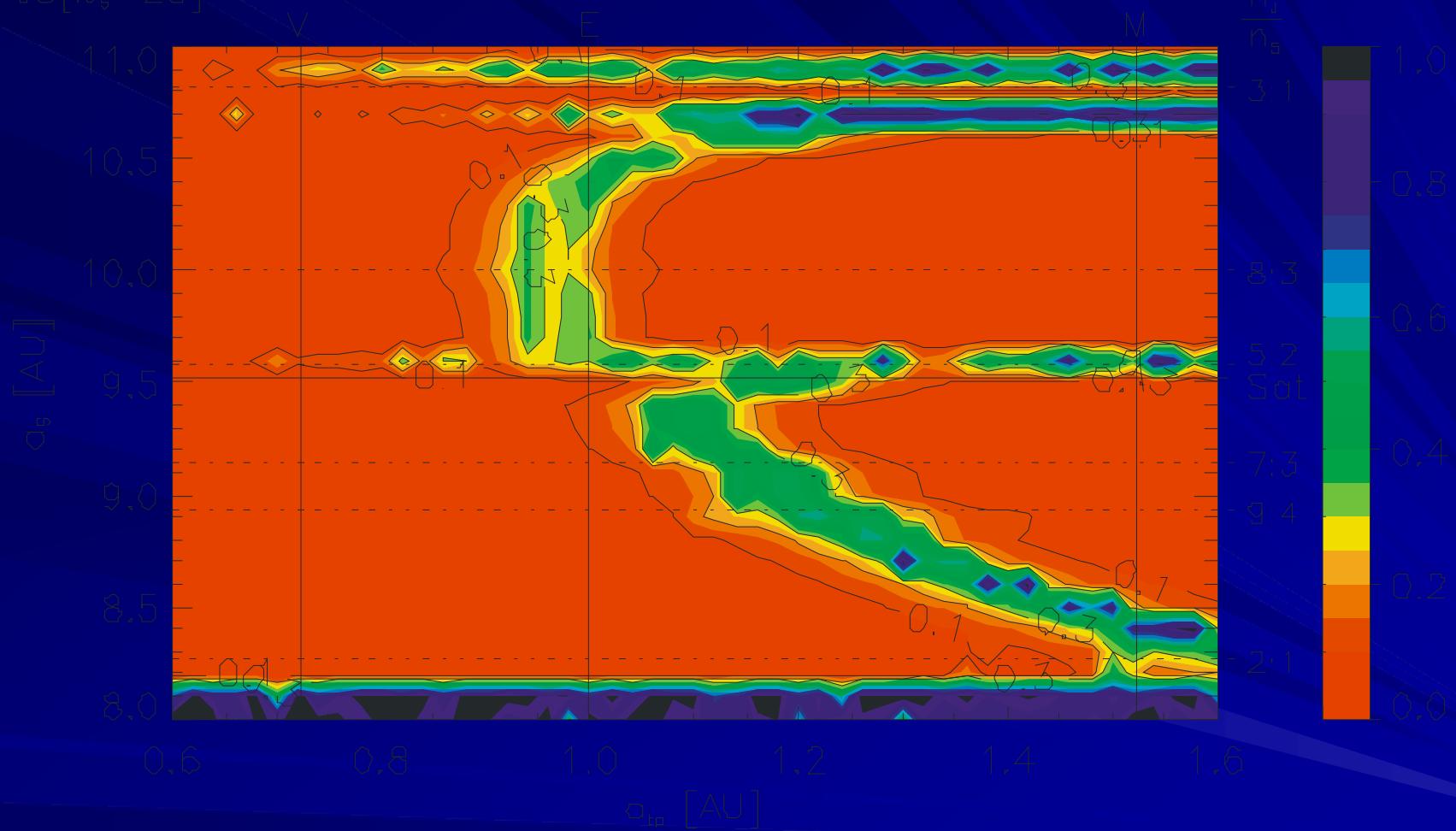
Jupiter -- Saturn



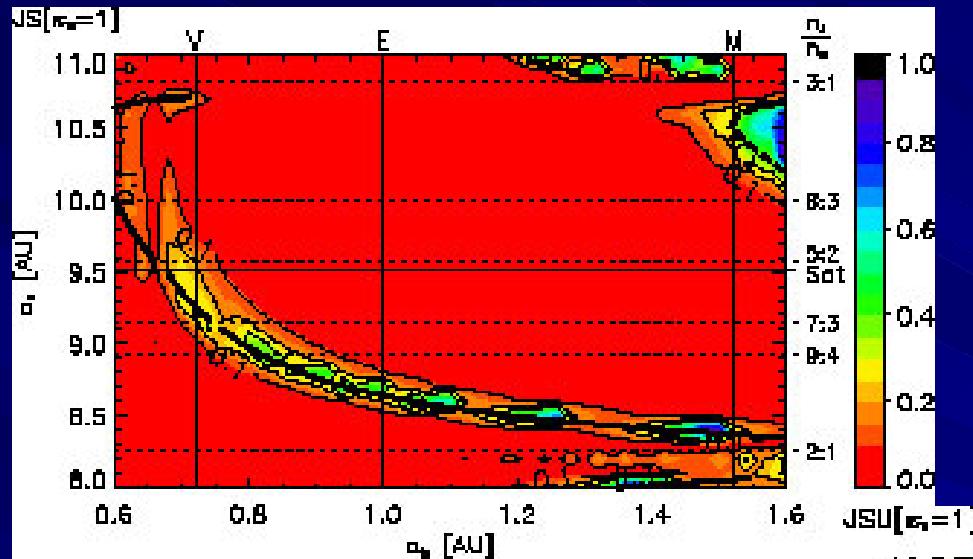




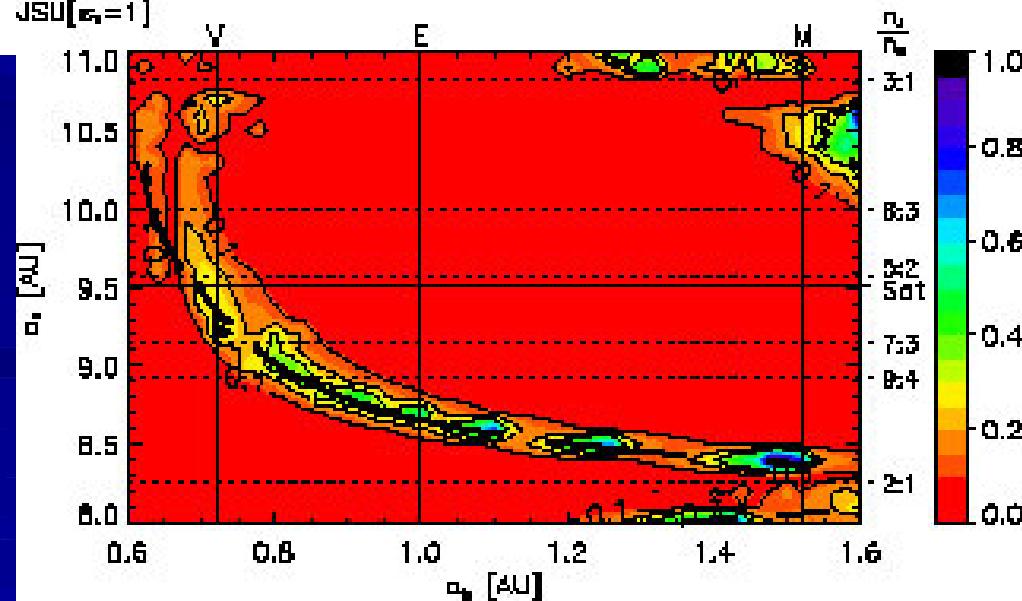
JS [$\kappa_s = 20$]

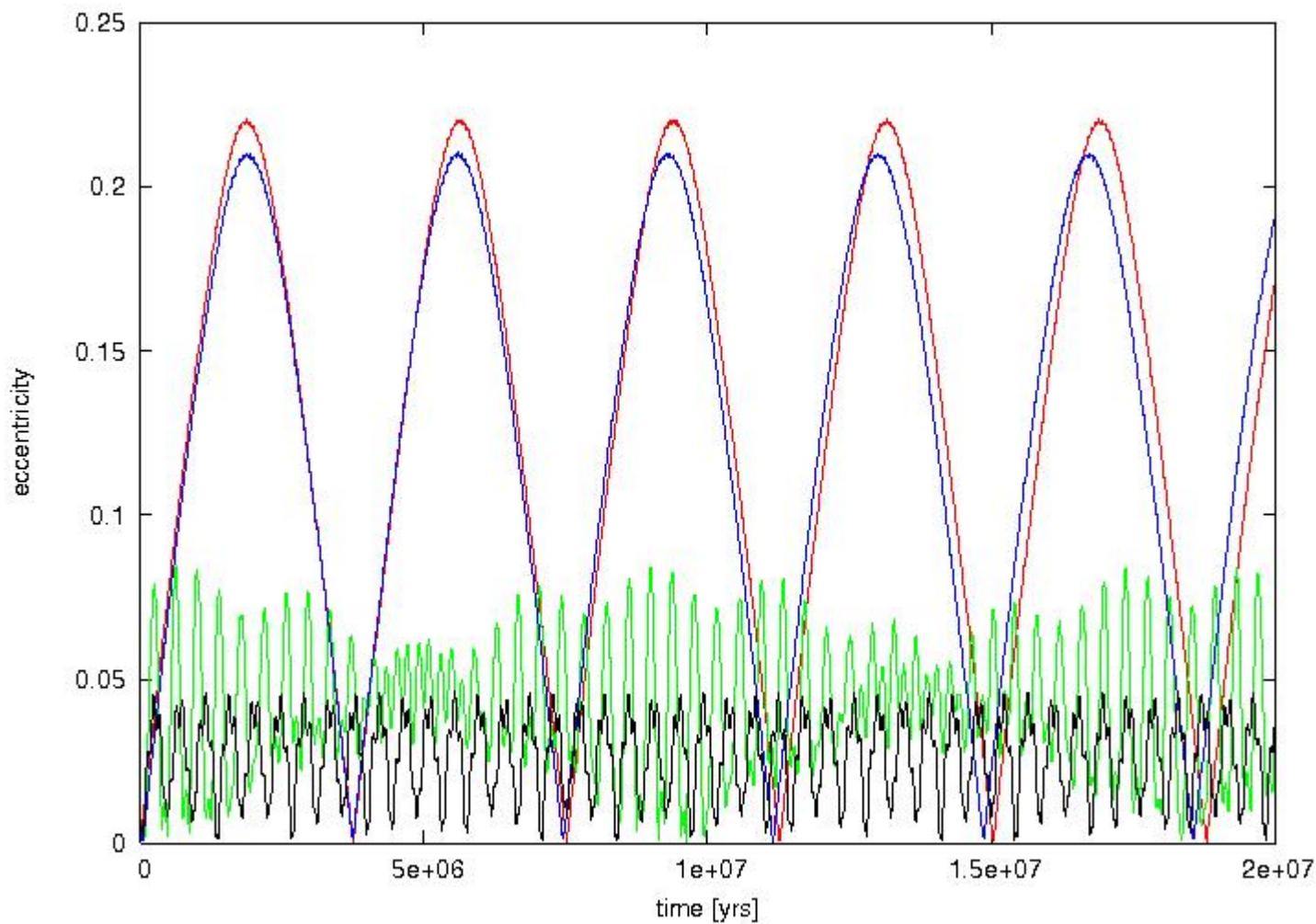


Influence of a third giant planet

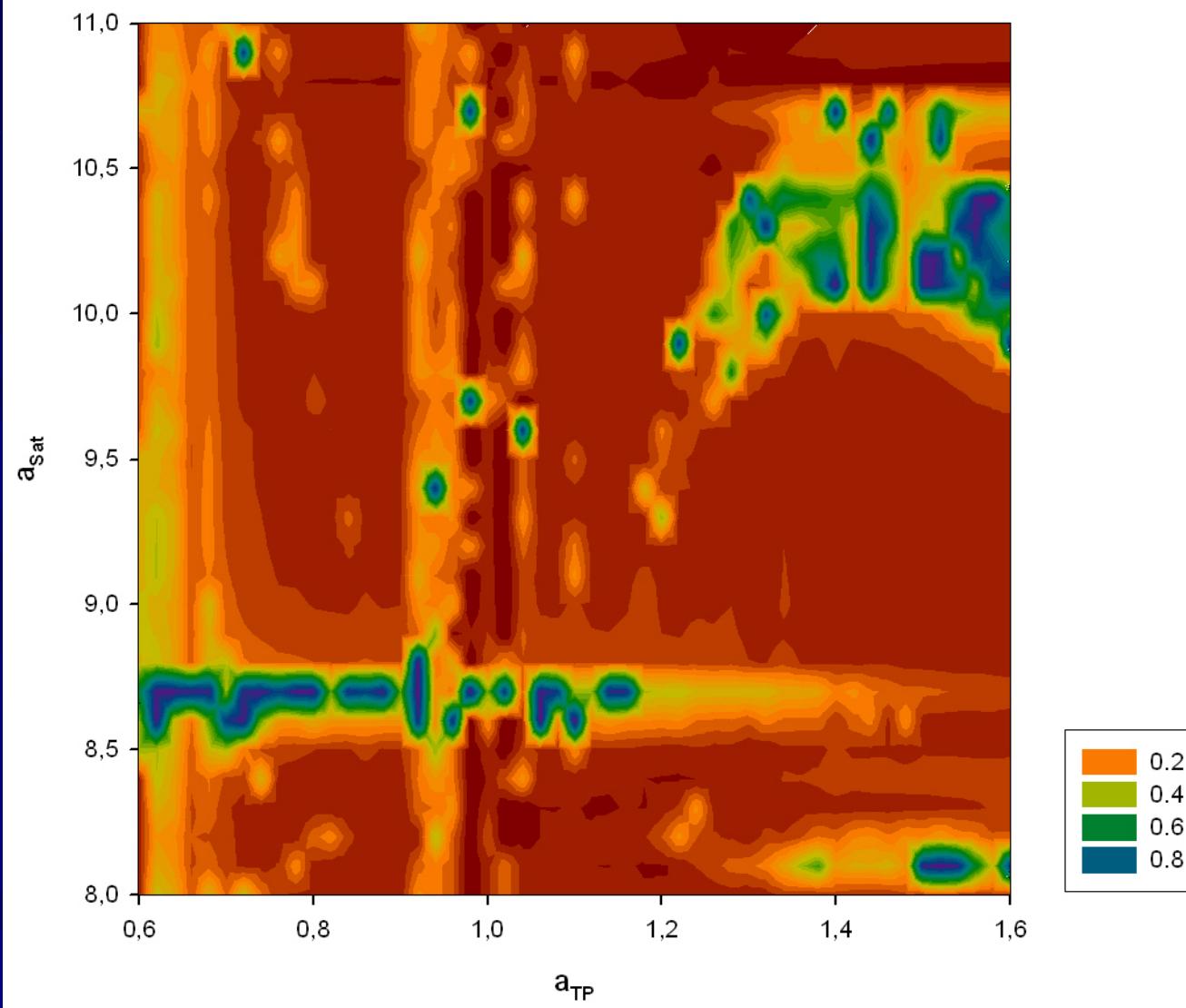


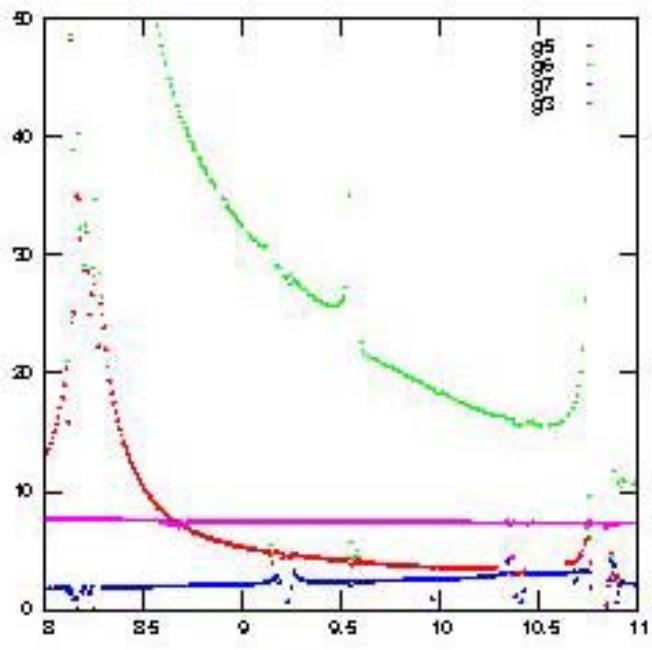
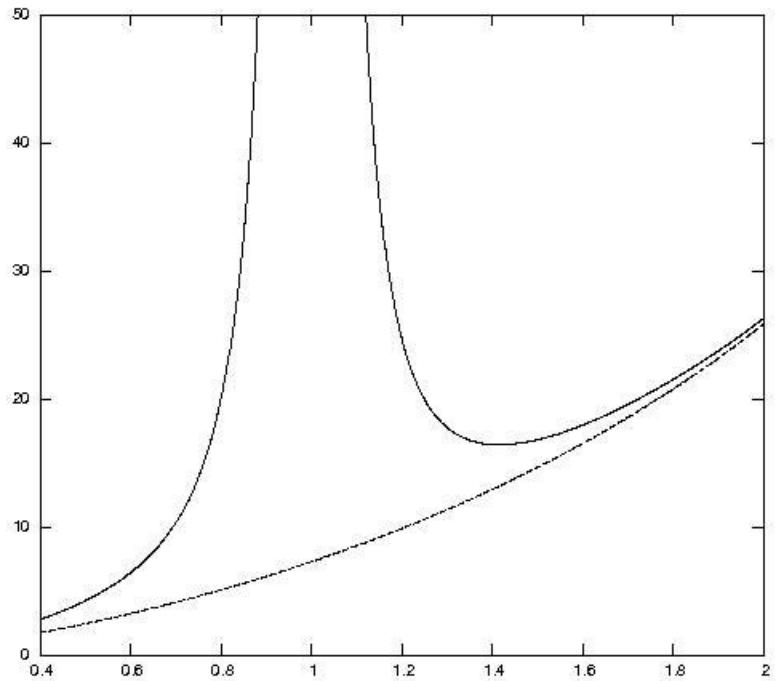
Jupiter -- Saturn

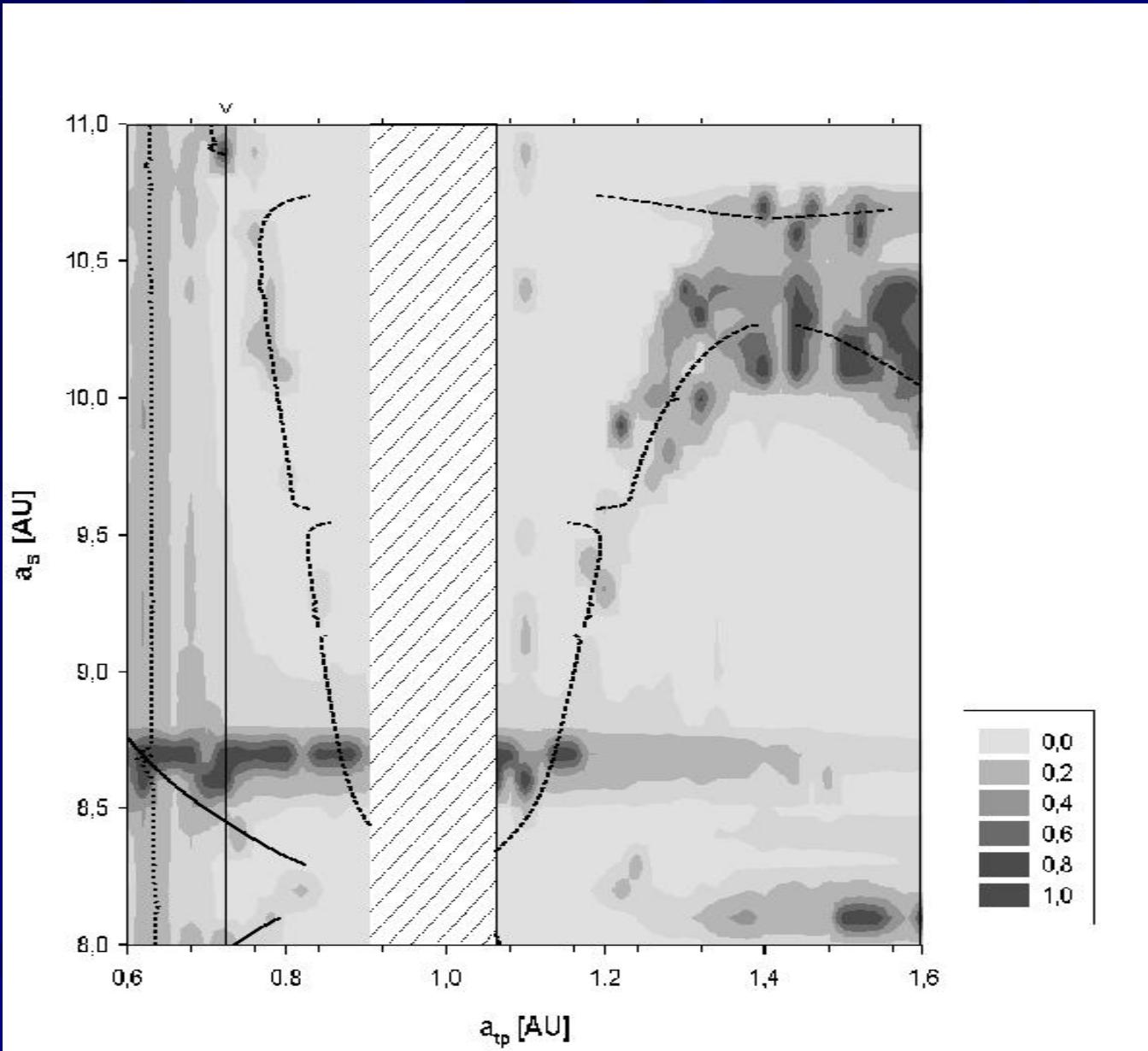




Sun - Jupiter - Saturn - Earth/Moon







Sun-Jupiter-Saturn-Uranus

Influence on an Earth-like planet at 1 AU

