

Some aspects of chaotic diffusion of Trojan asteroids

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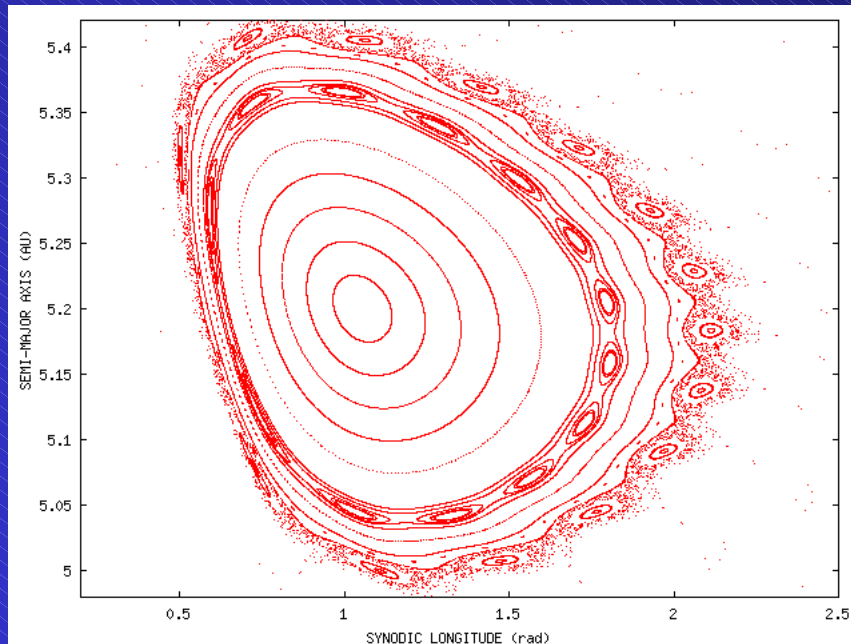
3rd Austrian Hungarian Workshop on Trojans

and Related Topics

May 13-15, 2002 Vienna

The phase space structure

Circular case: $e'=0$



**Inner chain of islands:
the 13:1 secondary
resonance**

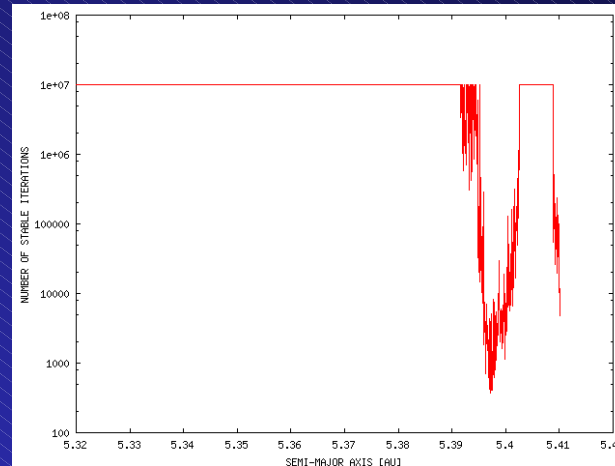
**Outer chain of islands:
the 14:1 secondary
resonance**

Boundary of the stability region

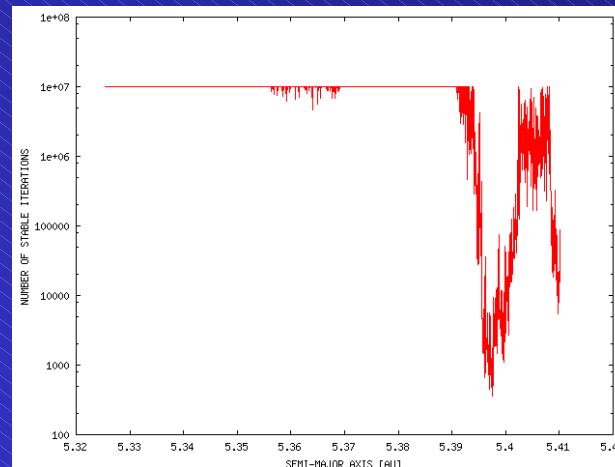
Elliptic case:

$$e' = 0.048$$

$$e = 0.048$$

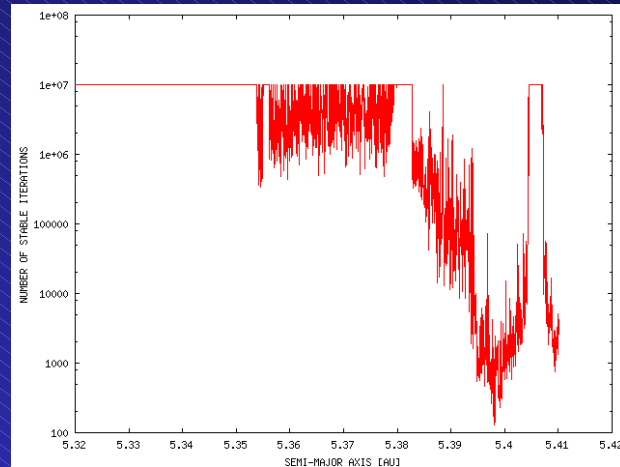


e', ϖ'
secularly
changing:
 $e = 0.048$

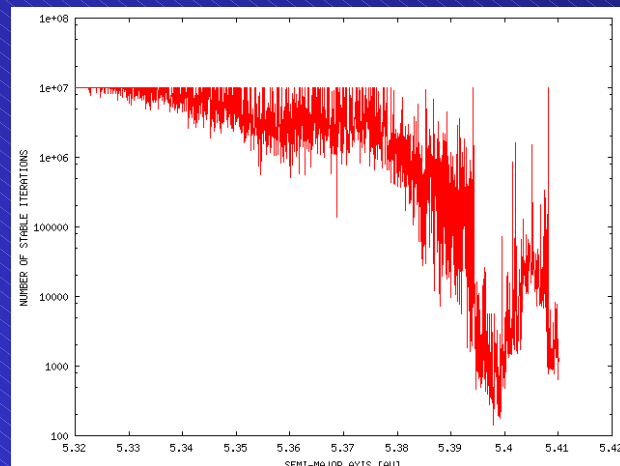


Boundary of the stability region

Elliptic case:
 $e' = 0.048$
 $e = 0.1$



e', ϖ' secularly
changing
 $e = 0.1$

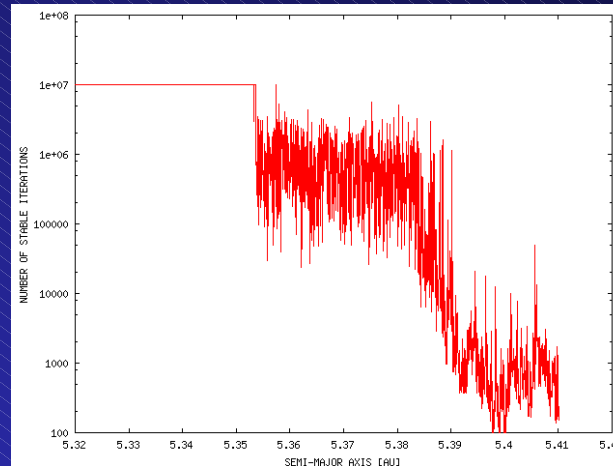


Boundary of the stability region

Elliptic case

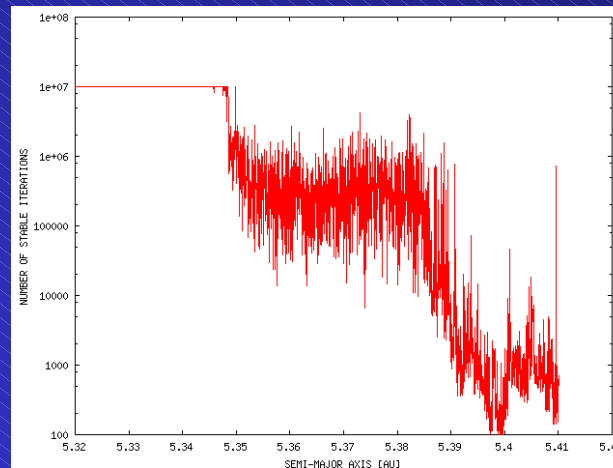
$e' = 0.048$

$e = 0.15$



e', ϖ' secularly
changing

$e = 0.15$

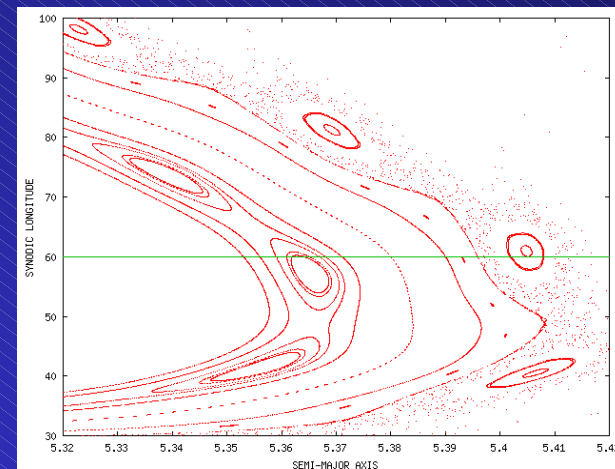
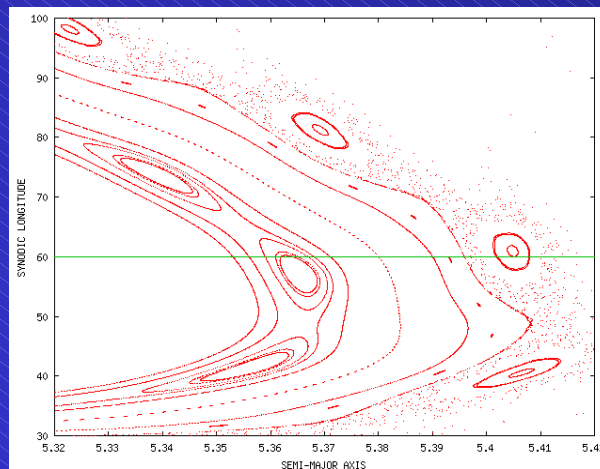
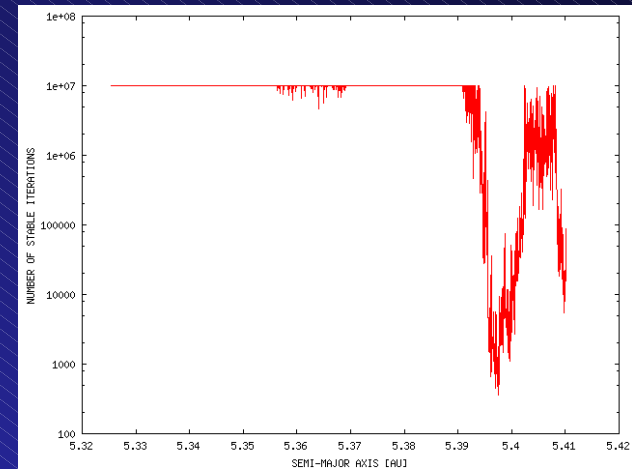
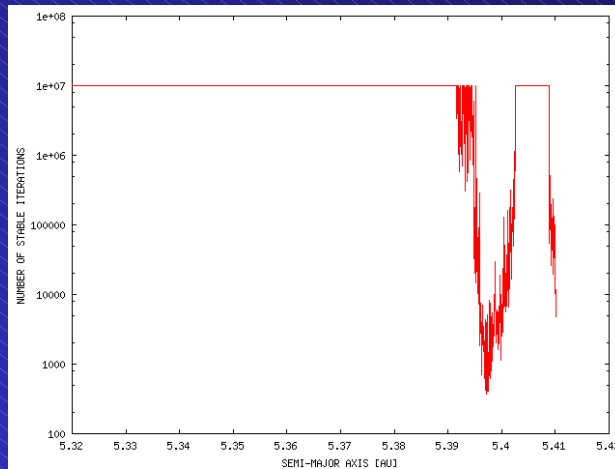


The role of the secondary resonances

pure elliptic case

secular case

$e = 0.048$

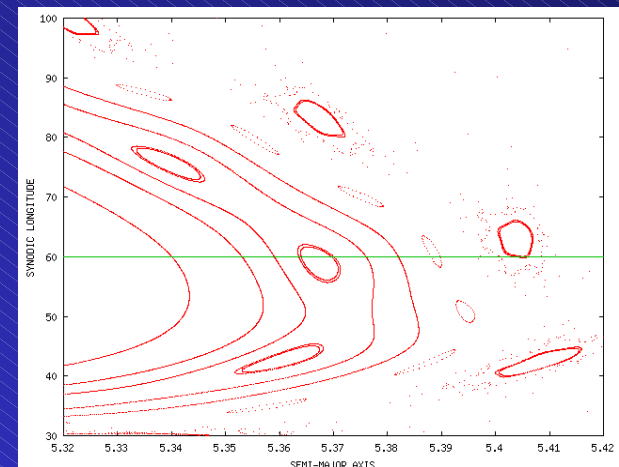
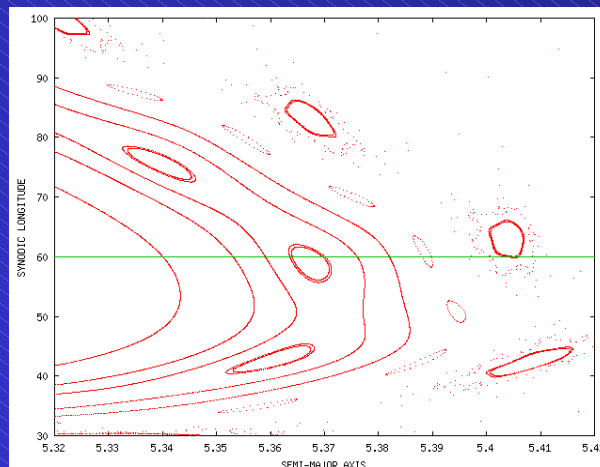
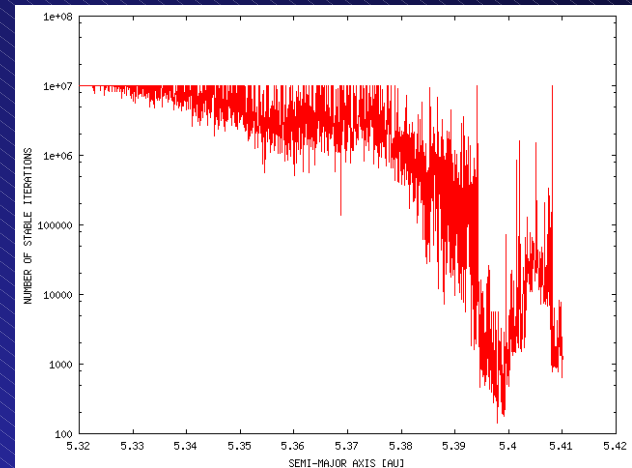
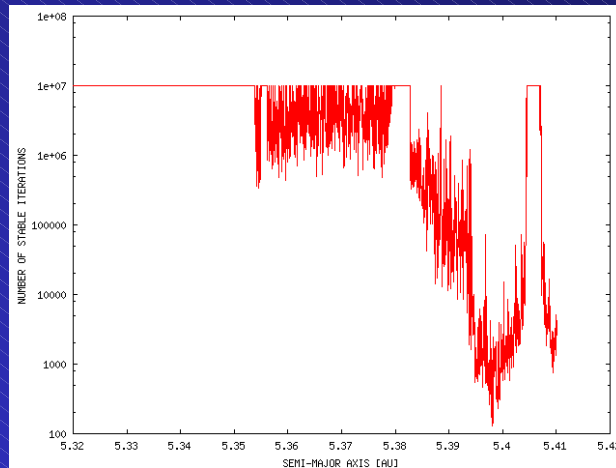


The role of the secondary resonances

pure elliptic case

secular case

$e = 0.1$

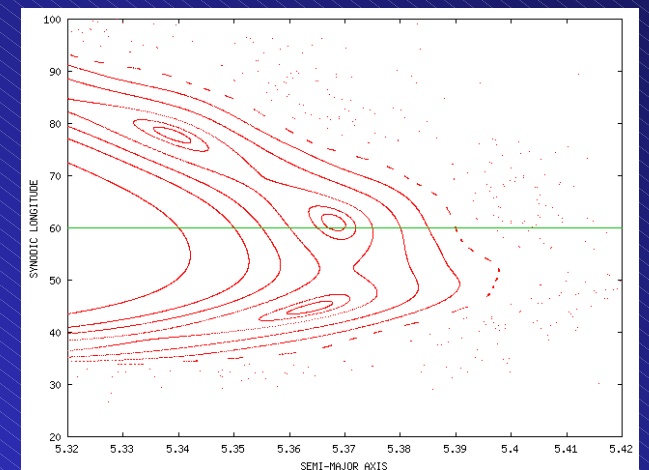
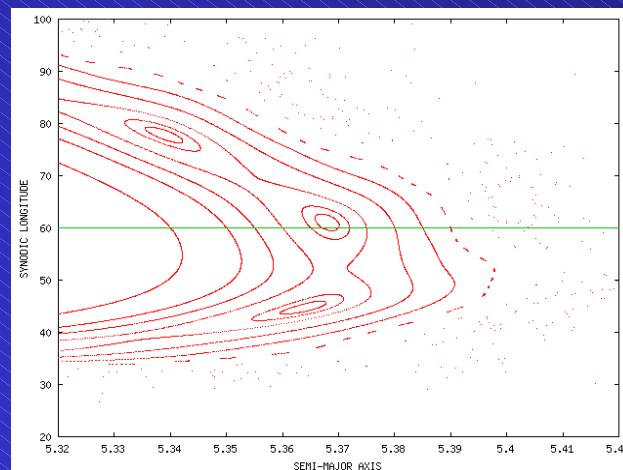
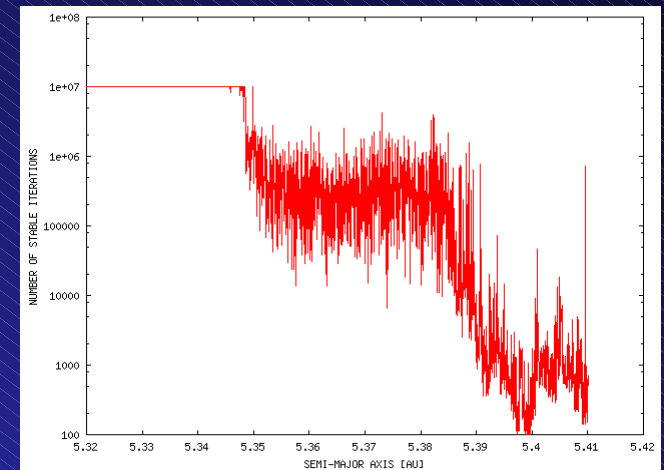
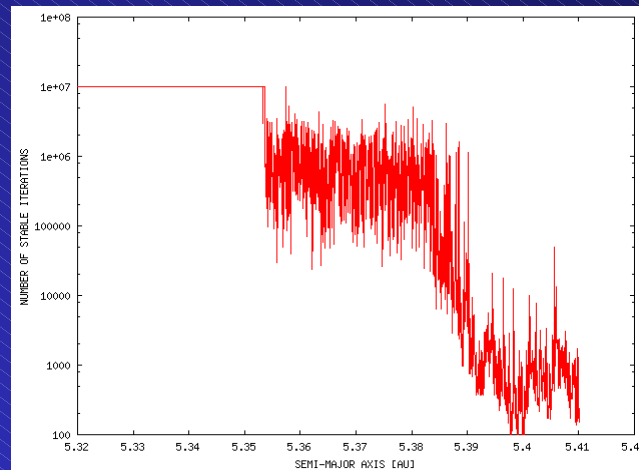


The role of the secondary resonances

pure elliptic case

secular case

$e = 0.15$



Property of orbits

Observation: orbits in the vicinity of the secondary resonances are chaotic in the elliptic case $e \neq 0$

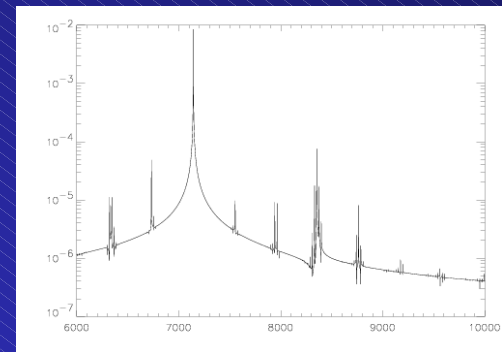
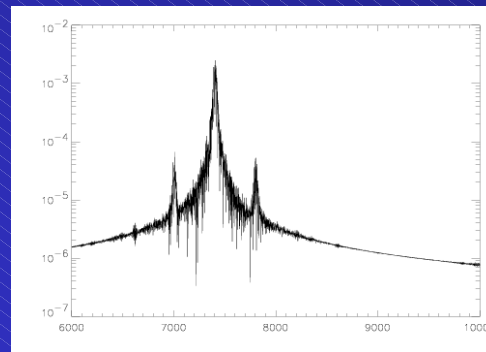
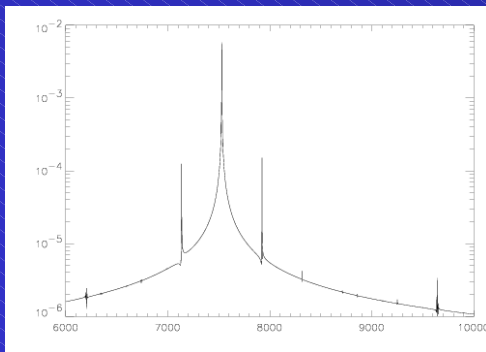
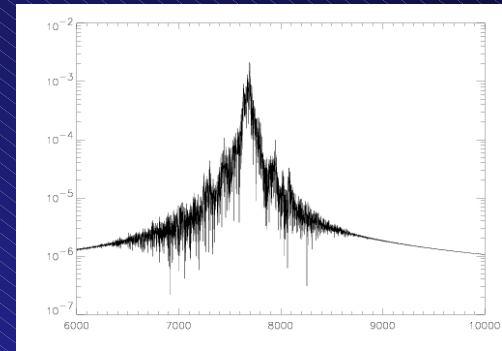
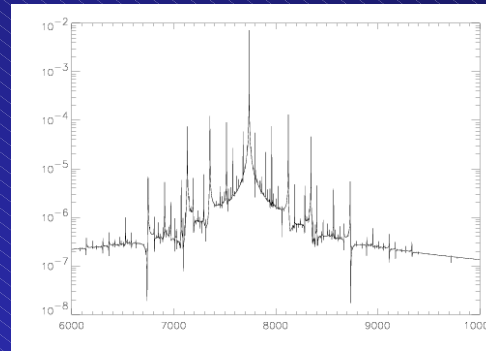
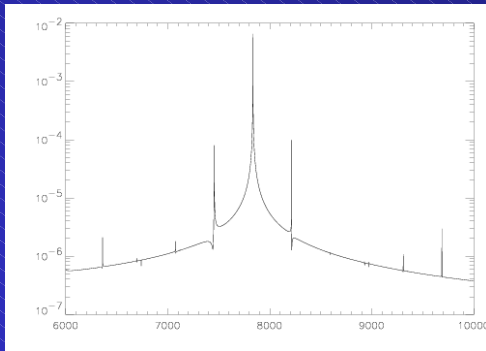
Proof: FFT power spectra

When e is small, orbits near these resonances are still stable (but chaotic), by switching on the planetary perturbations these regions become more unstable

By increasing e these regions become unstable even in the elliptic case, the presence of secular perturbations do not influence very much the stability

FFT Power Spectra

$e = 0.048$; $a = 5.346, 5.357, 5.367, 5.378, 5.388, 5.40$ (AU)



Conclusions:

The regions of the phase space, where in the circular problem ($e'=0$) the islands of secondary resonances take place, become chaotic by increasing e' .

For small e' these chaotic regions are stable, by switching on the secular perturbations these regions become unstable in a 10-100 million years timescale.

For larger e' , these regions are unstable, by adding the secular perturbations, the stability region does not change essentially.

The secondary resonances may play a not negligible role in the destabilization of Jupiter's Trojans