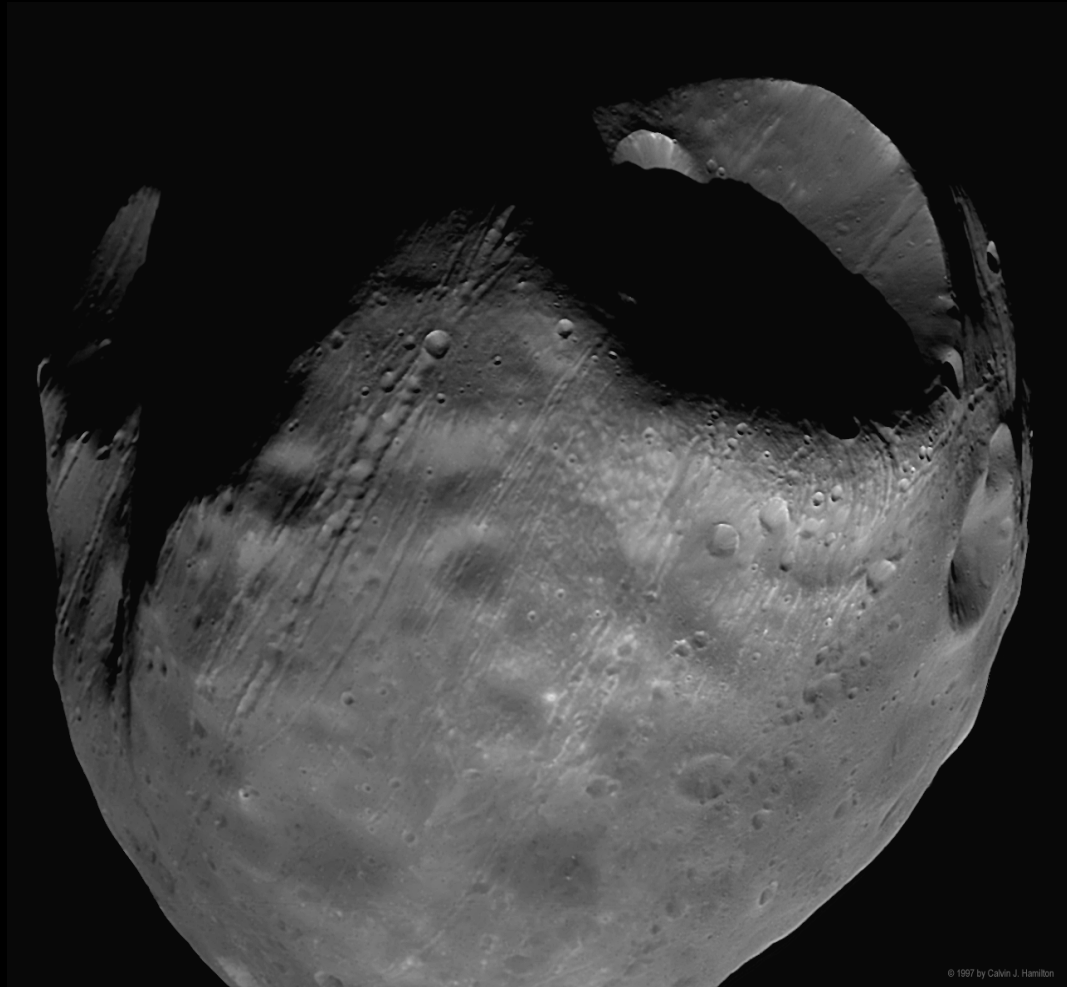


# The Rotation of Natural Satellites



3<sup>th</sup> Austrian-Hungarian Workshop  
on Trojans and related Topics

# Spin influenced by

- Wobble decay
- Tides
- Spin orbit coupling

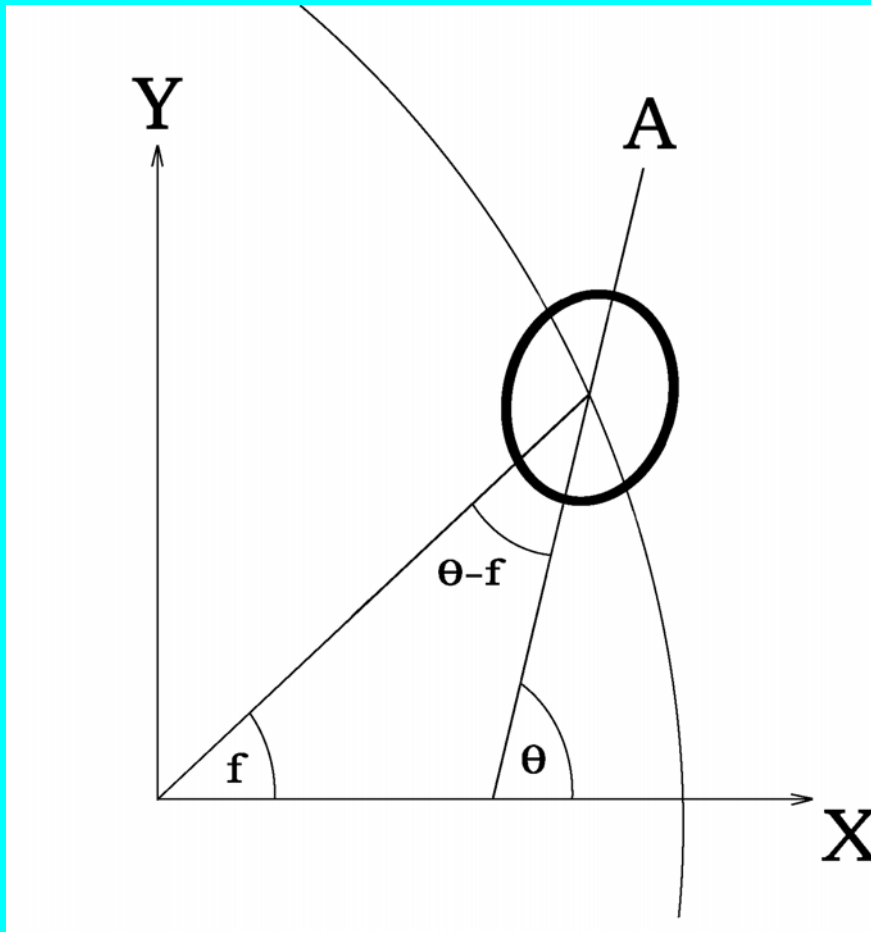
# Wobble decay:

- Periodic distortion of the body  
rotation of the satellite  $\rightarrow$  principal axis  
rotation
- Time scale: small

# Tides:

- spin-angular momentum  $\rightarrow$  orbit
- gravitational field distorts the satellite  
phase lag  $\delta$  in the response of grav. field
- Ocean tides
- solid body tides

# Definitions



Edge-on-view of the satellites' orbit

$$\frac{d^2\theta}{dt^2} + \frac{3}{2} \frac{B-A}{C} n^2 \left(\frac{a}{r}\right)^3 \sin(2(\theta - f)) = W$$

$W$  tidal torque

$a$  orbital semimajor axis

$r$  distance between the planet and the satellite

$A, B, C$  principal axes

$n$  orbital mean motion

$f$  true anomaly

# Circular orbit

- Angular Momentum  
satellites spin  $\rightarrow$  orbit  
synchronous rotation

# Non circular orbit

- Tidal Torque  $W$

Time rate of changing the spin angular velocity

$$W = \frac{d\omega}{dt} = -\frac{3}{2} \frac{k_2 GM^2 R^5}{r^6} \sin(2\delta) = -\frac{45\rho R^4 n^4}{38\mu Q} \left(\frac{a}{r}\right)^6$$

$$\frac{d\theta}{dt} = \omega$$

M

R

r

$\mu$

$\rho$

a

Q

spin angular velocity

planetary mass

mean equatorial radius

distance between the planet and the satellite

rigidity of the satellite

mass density

orbital semimajor axis

specific dissipation function

$$\frac{1}{Q} = \frac{1}{2\pi E^*} \oint \frac{dE}{dt} dt \quad \frac{1}{Q} \approx 2\delta$$

$E^*$  peak energy stored in an oscillating system

$\oint \frac{dE}{dt} dt$  energy dissipated over a cycle of oscillation

# Spin orbit coupling:

- End point of tidal evolution

$$\langle W \rangle = 0$$

- non spherical shaped satellites

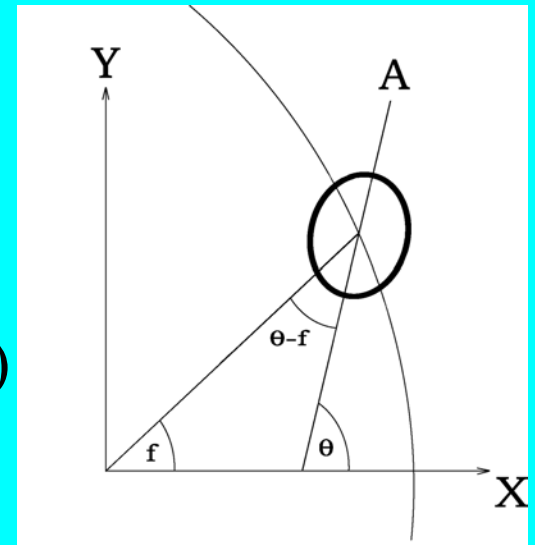
— > resonant torque (caused by spin orbit coupling) possible

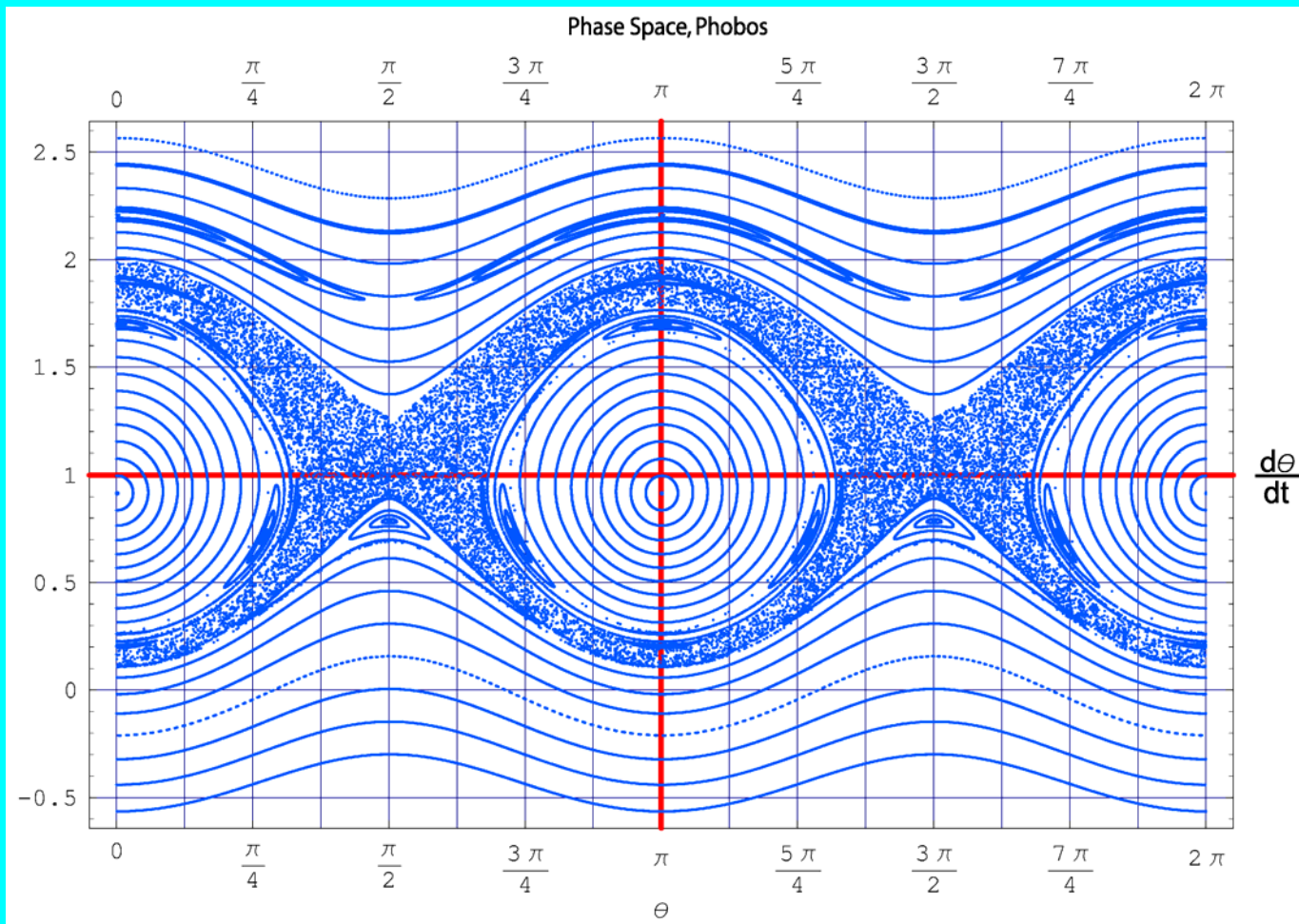
$$\frac{d\theta}{dt} = pn \quad n = \frac{2\pi}{T}$$



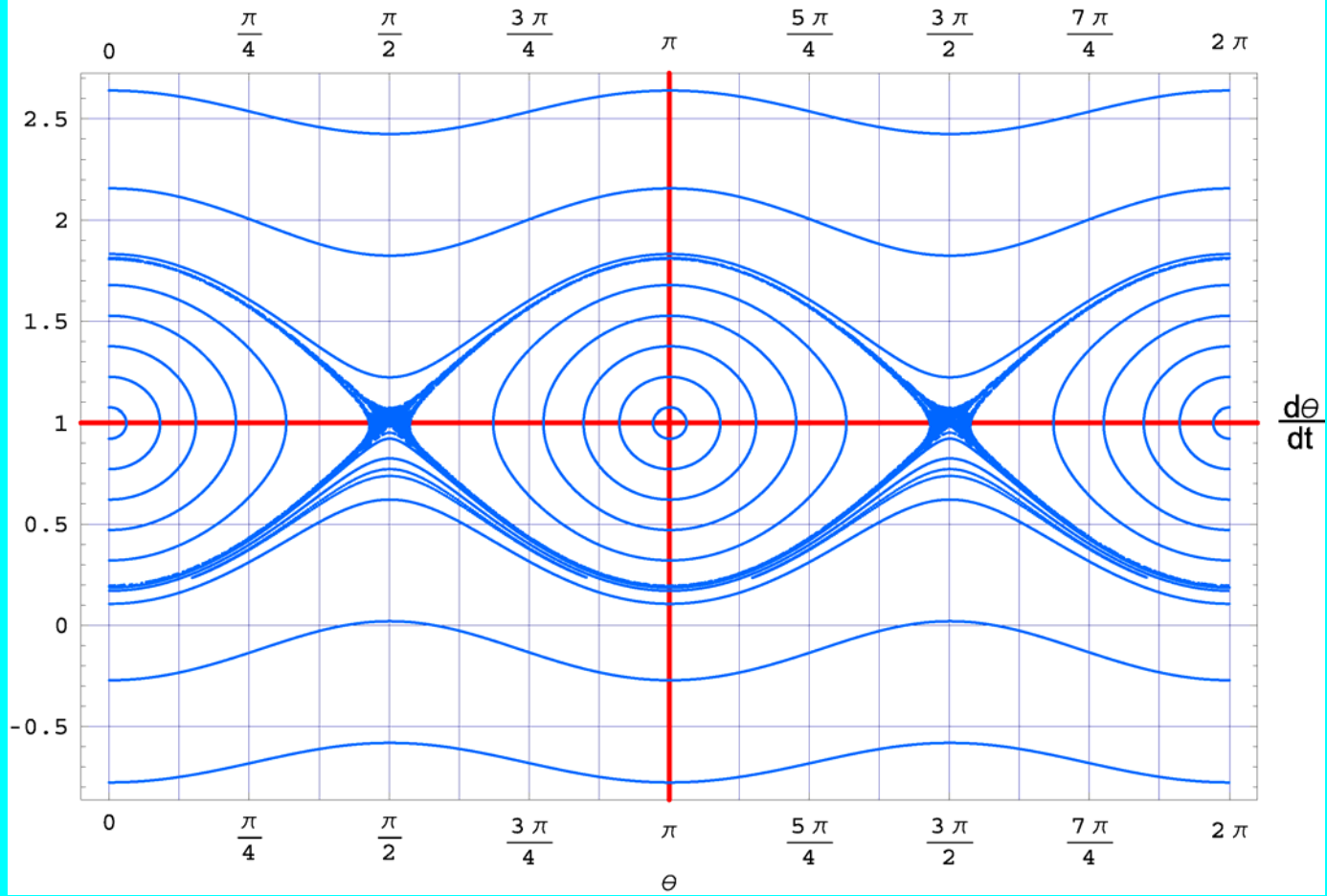
# Phobos - Deimos

- Equation  $\frac{d^2\theta}{dt^2} + \frac{3}{2} \frac{B-A}{C} n^2 \left(\frac{a}{r}\right)^3 \sin(2(\theta - f)) = W$  for  $W=0$  was solved.
- Runge Kutta 4<sup>th</sup> order
- Program generator for Mathematica execution — > Plots: spin angular velocity  $\frac{d\theta}{dt}$  vs.  $\theta$  at perihelion
- For some special initial conditions:  
limb profiles at perihelion  
frequency analysis (FFT with  $2^{20}$  Points)  
Lyapunov Exponents

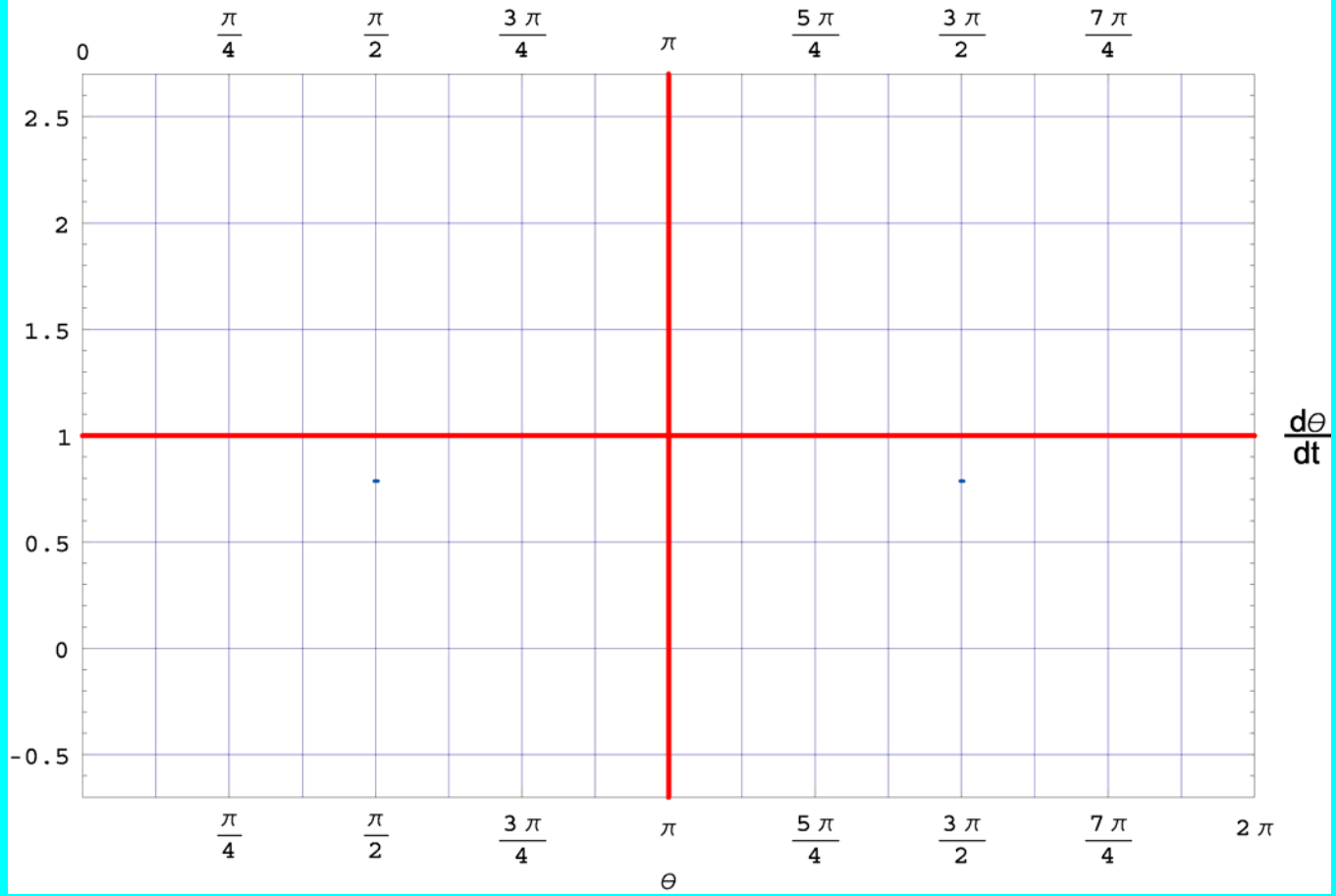


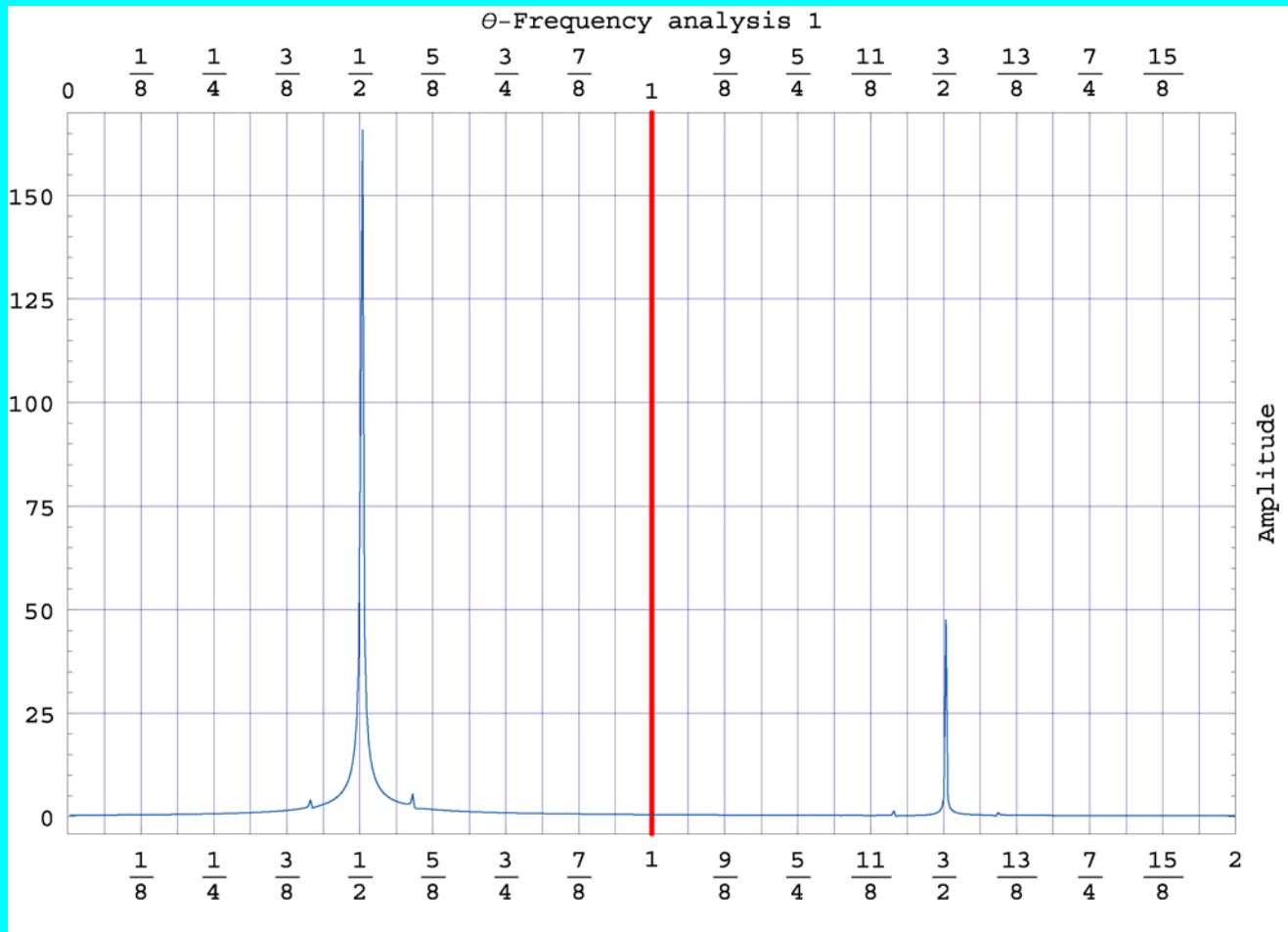


Phase Space, Deimos

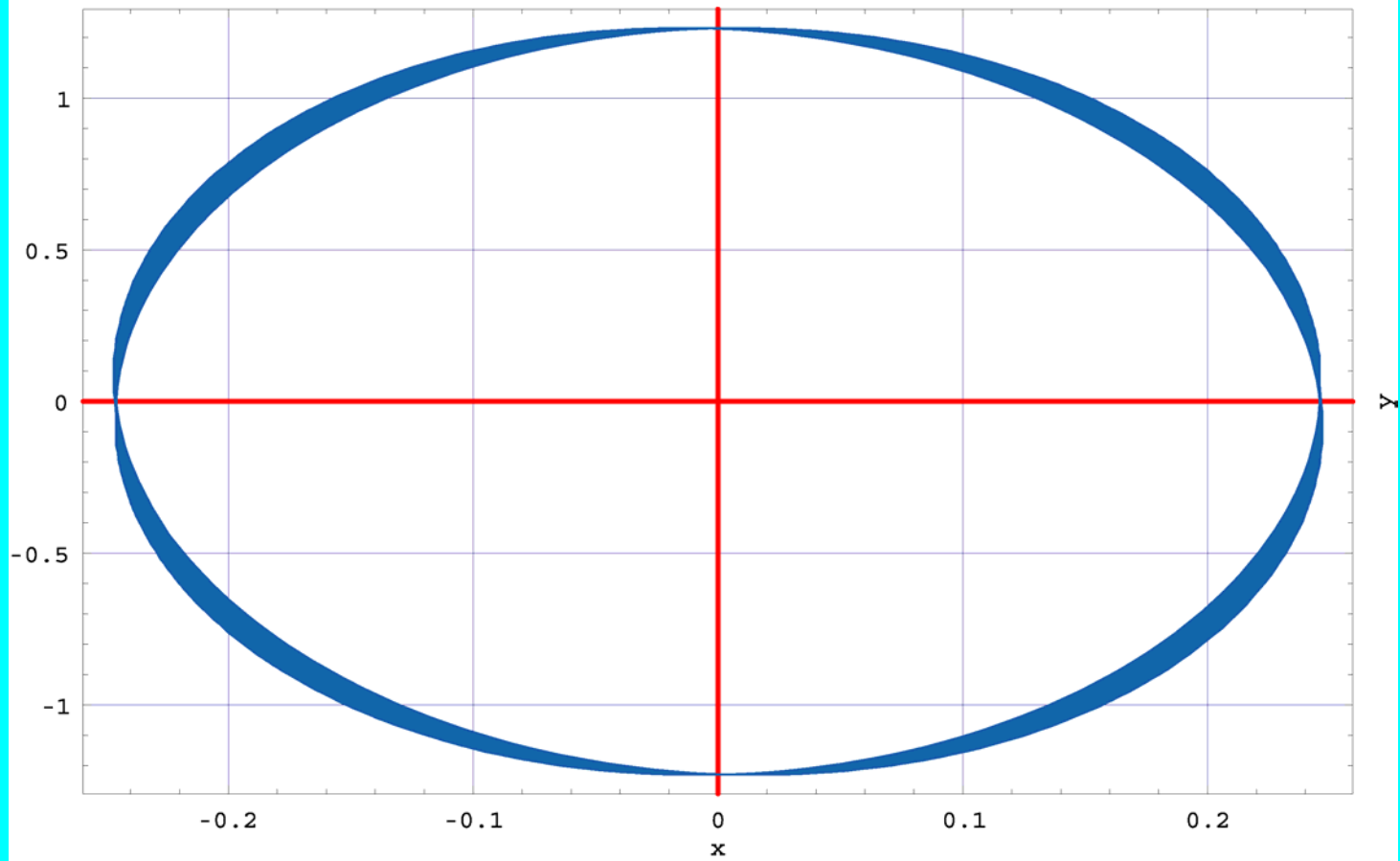


Phase Space 1, Phobos

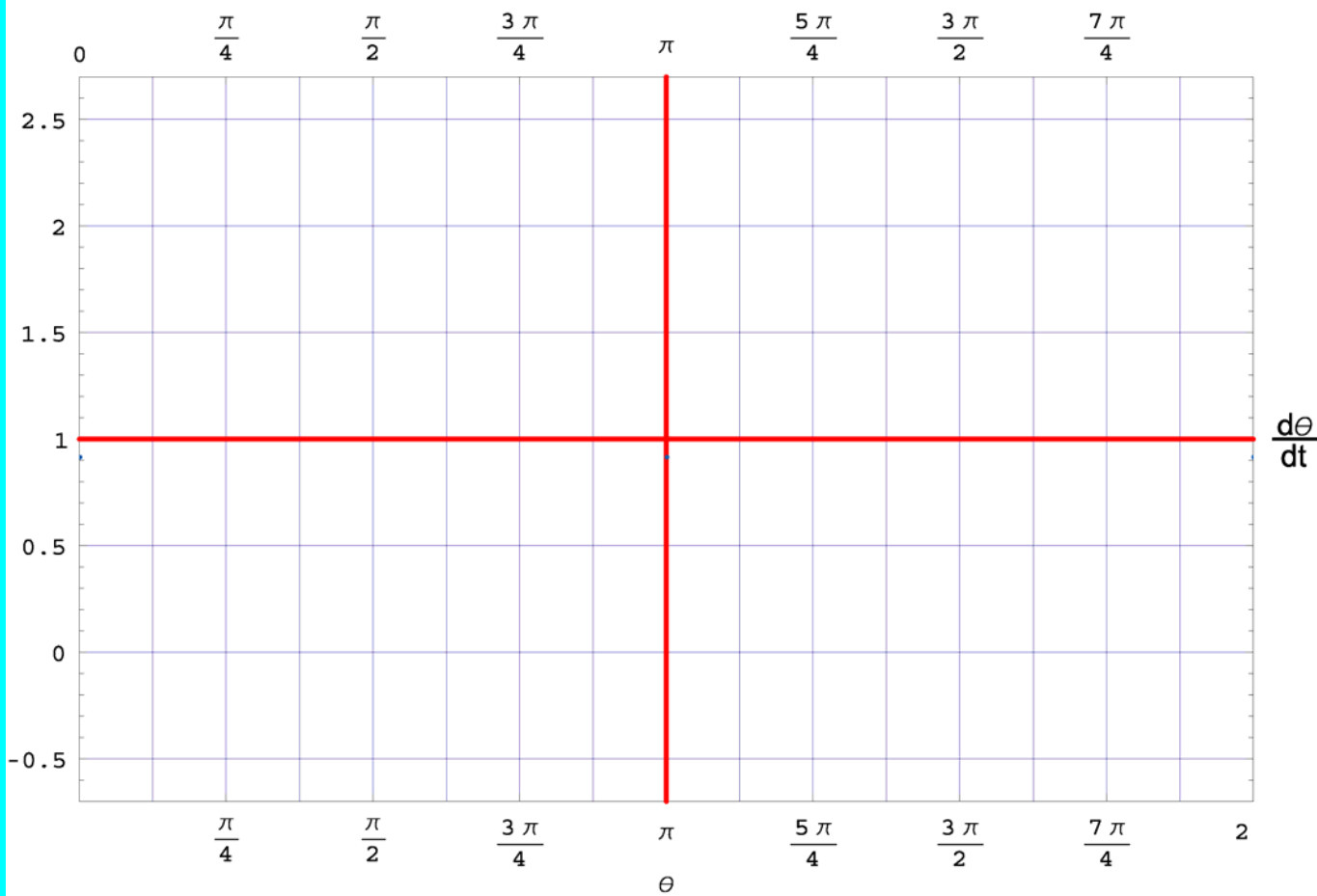


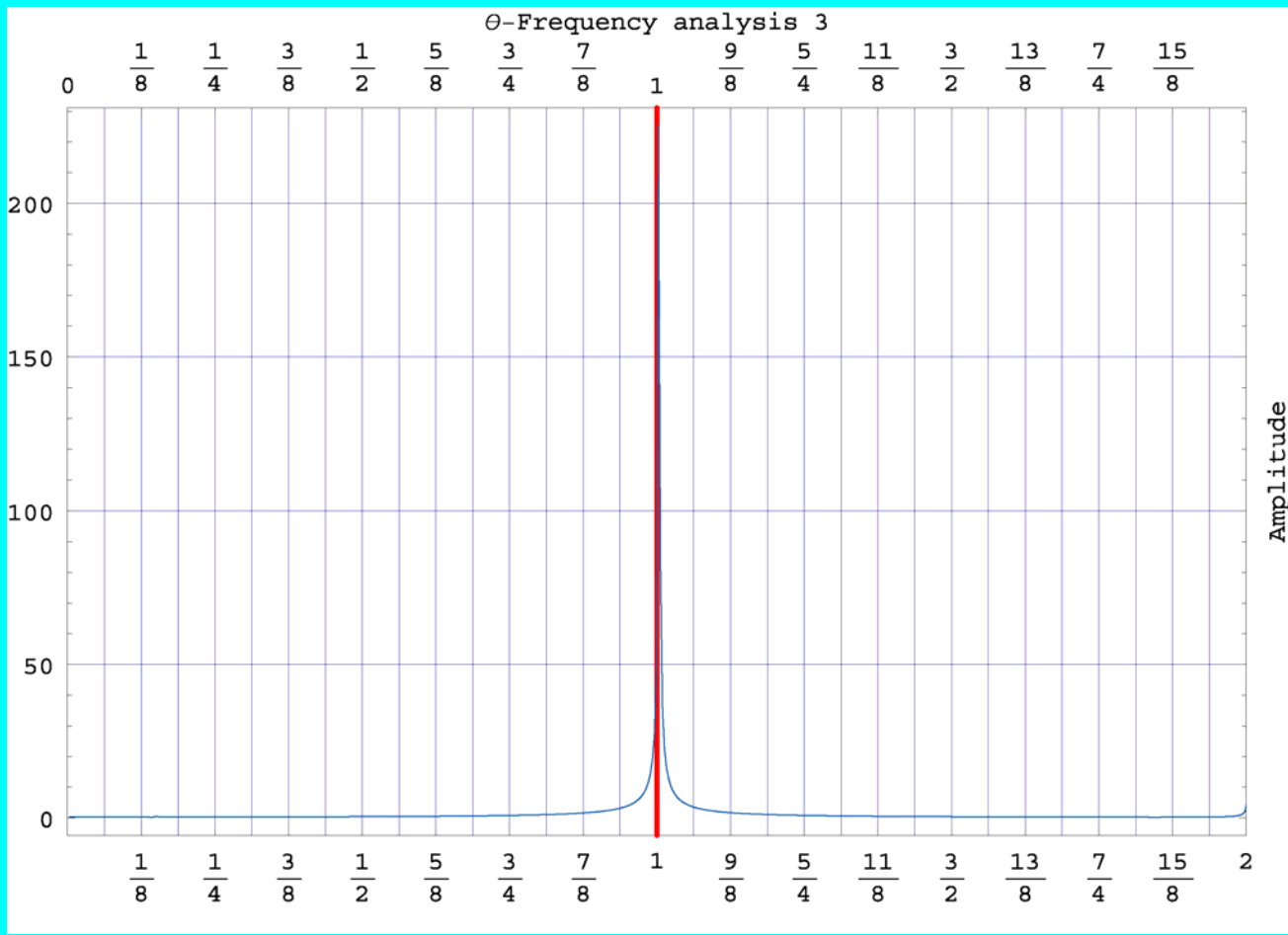


Phobos - Limb profile 1



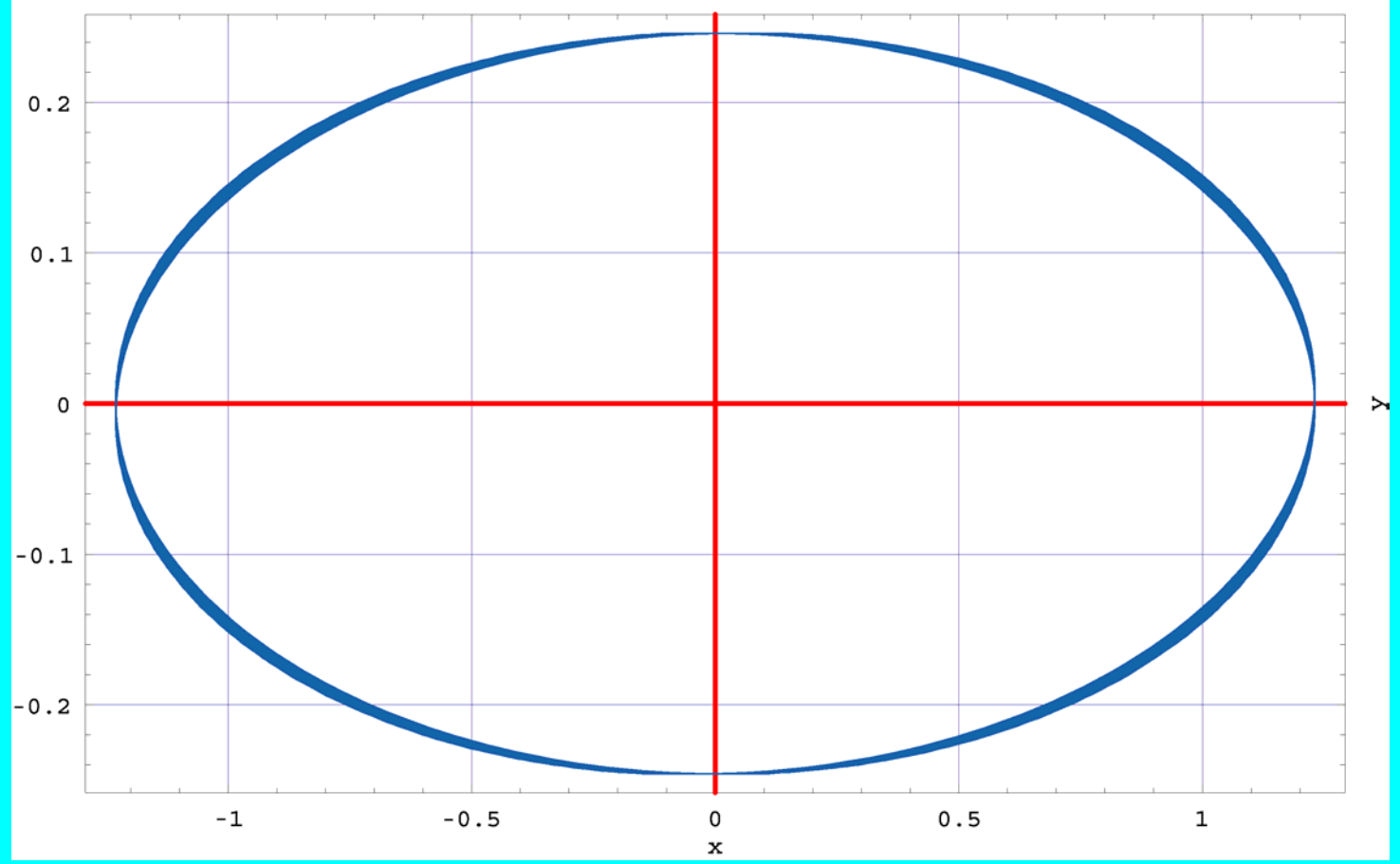
Phase Space 3, Phobos



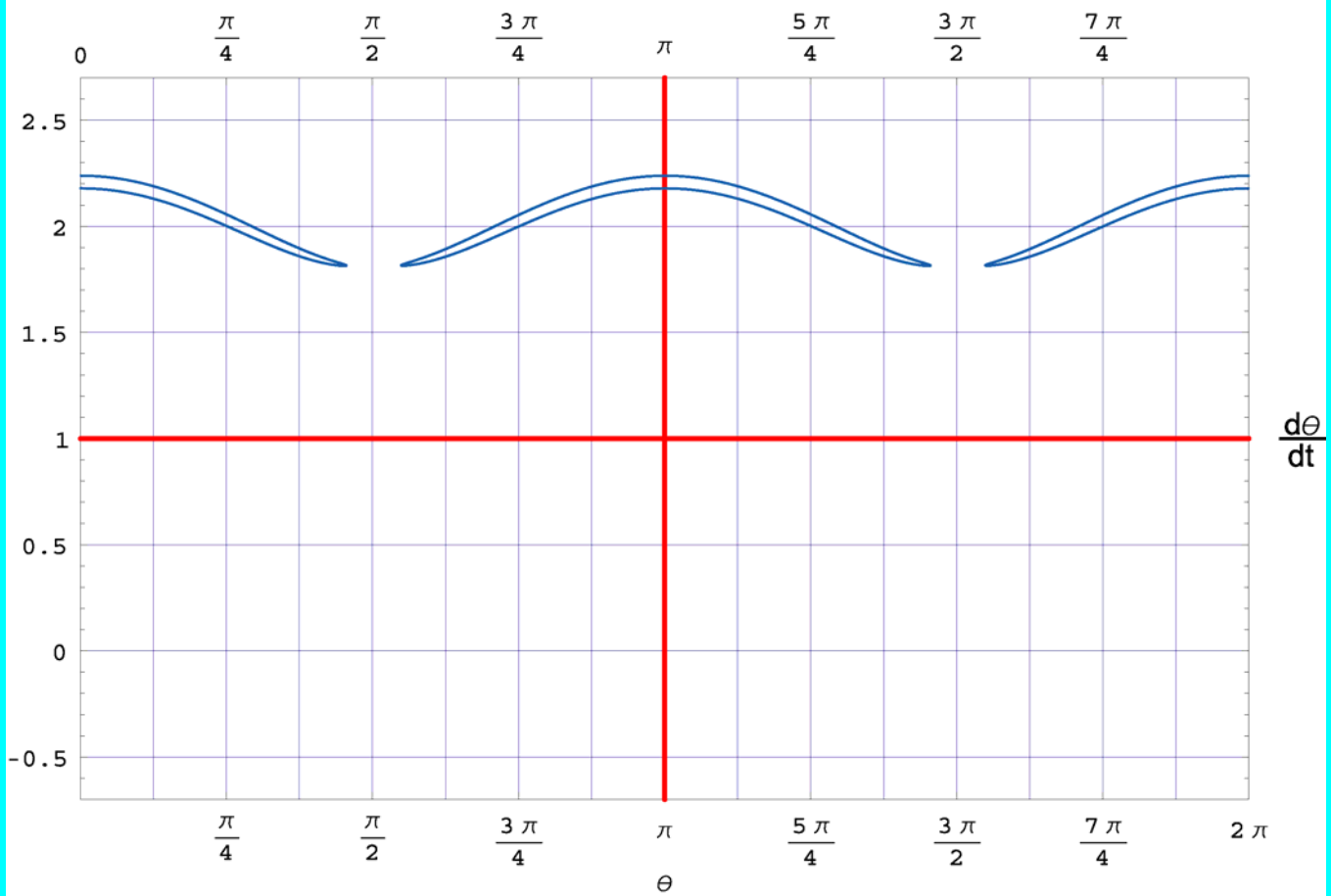


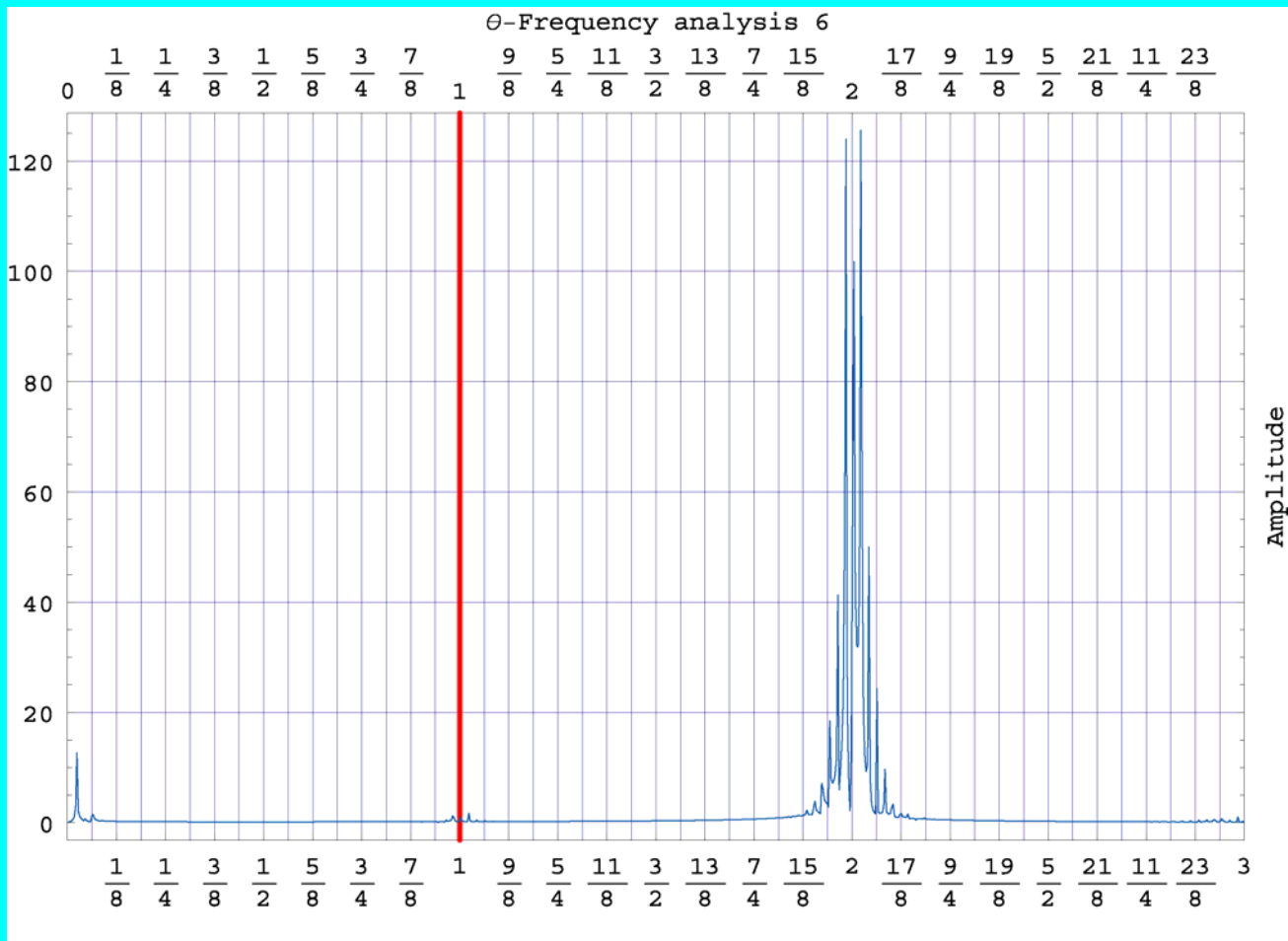


Phobos - Limb profile 3

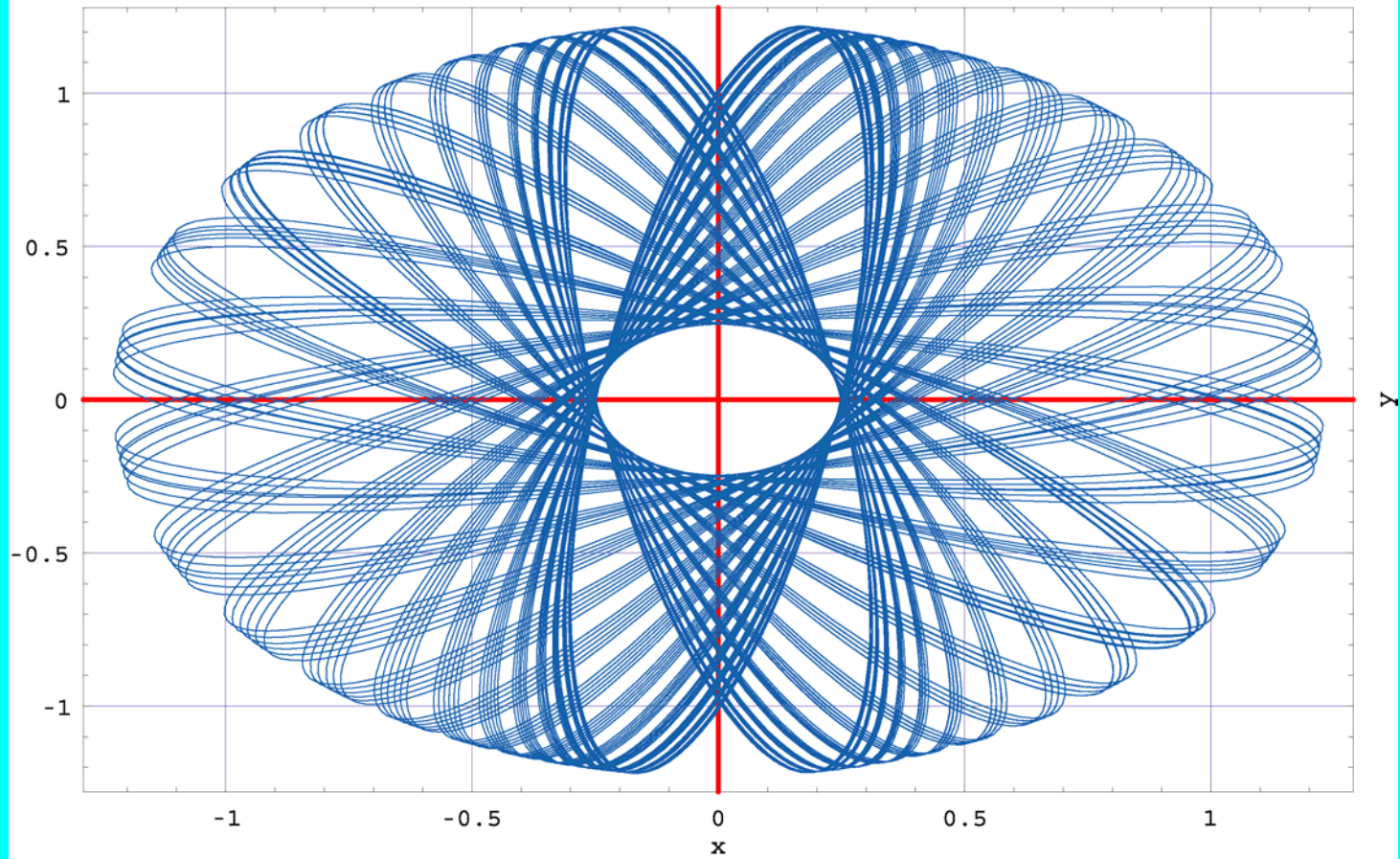


Phase Space 6, Phobos

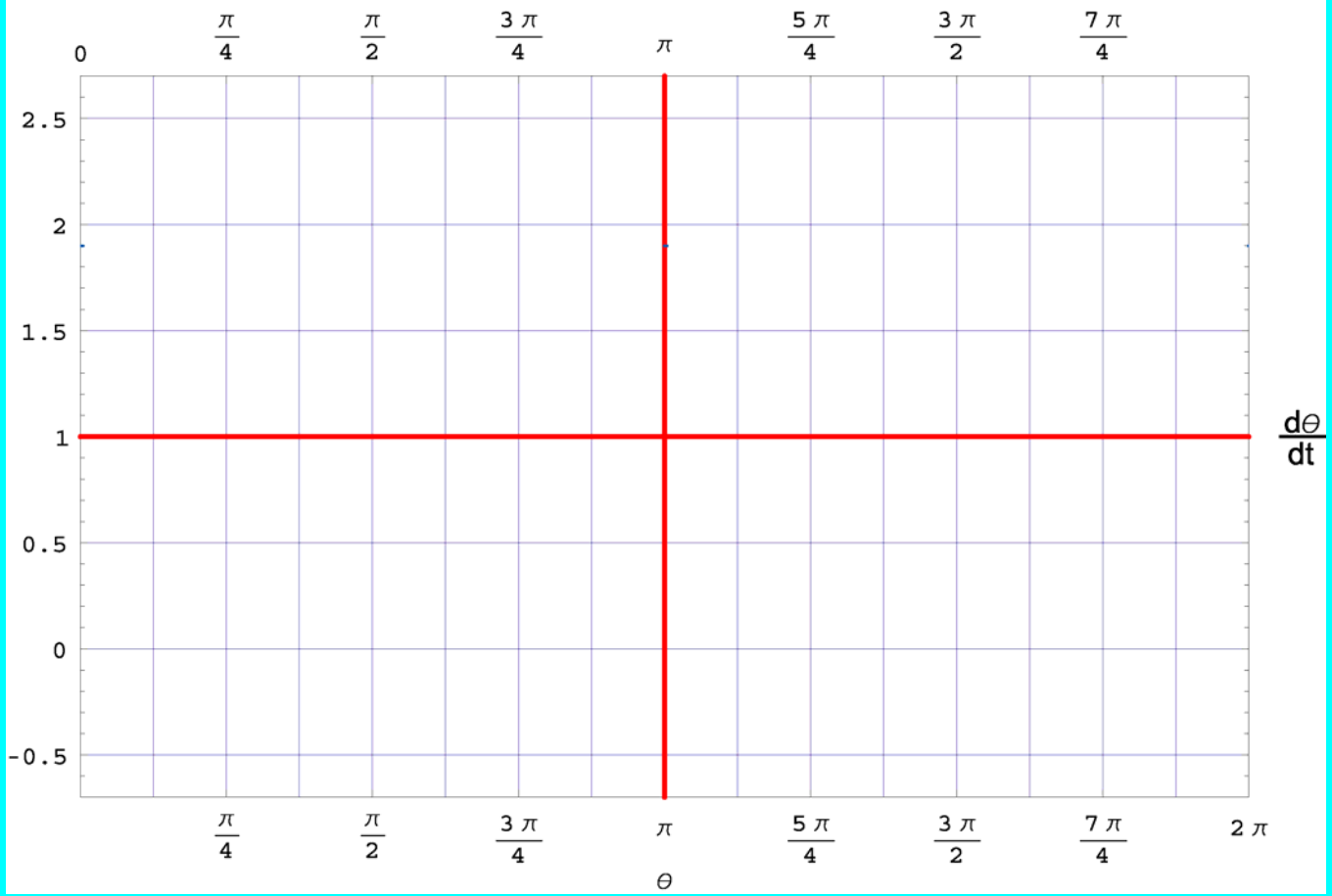


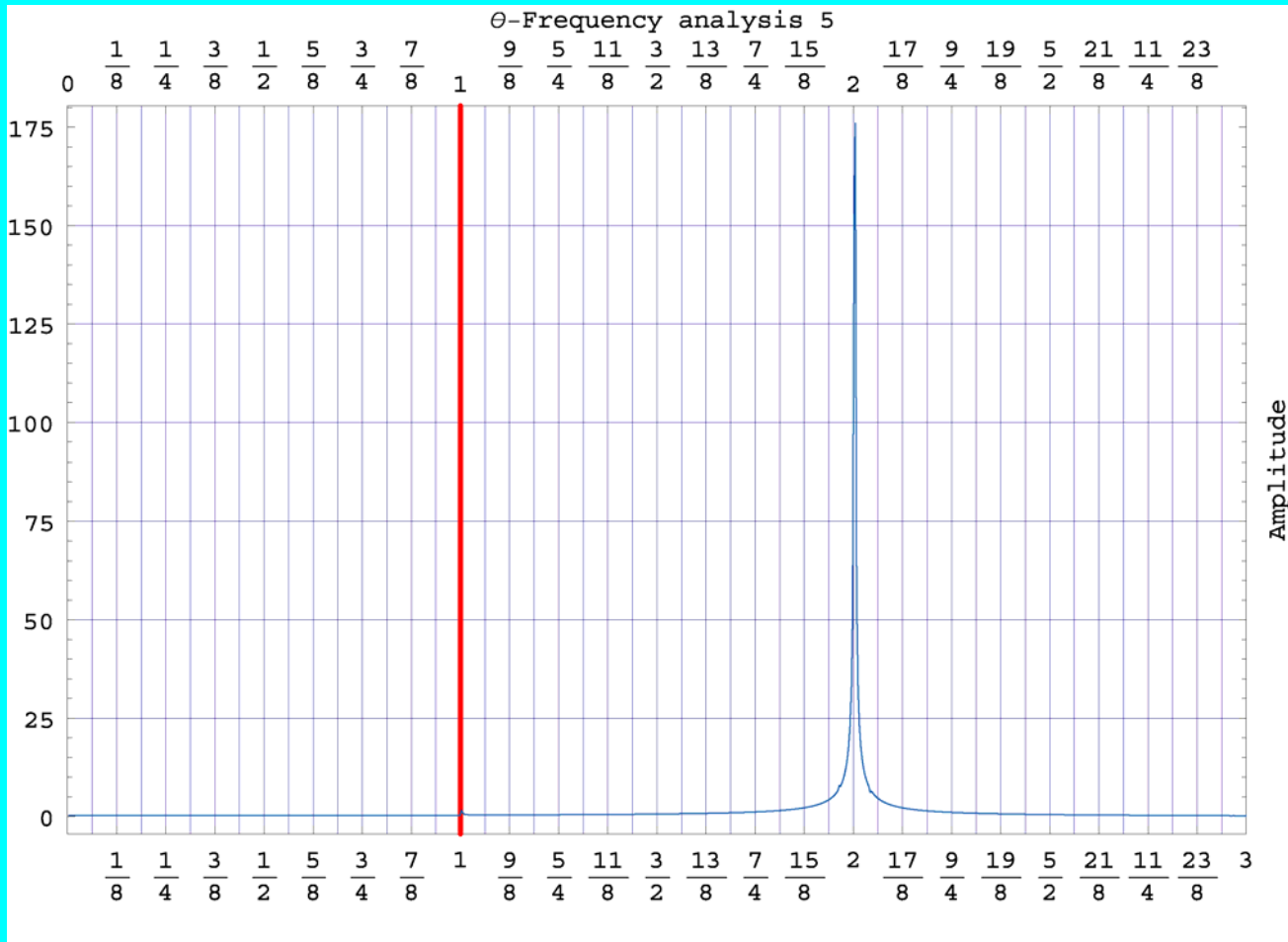


Phobos - Limb profile 6

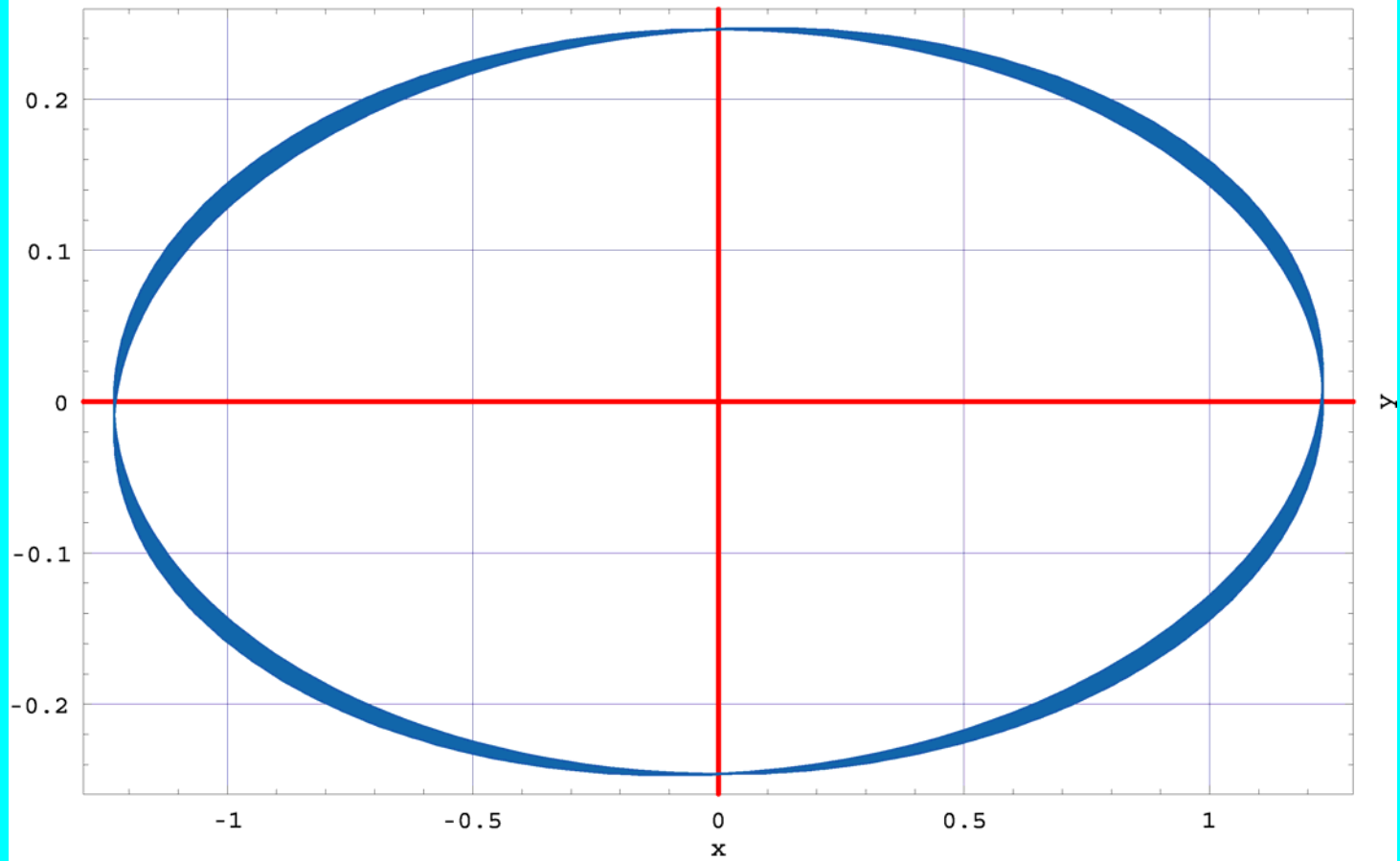


Phase Space 5, Phobos

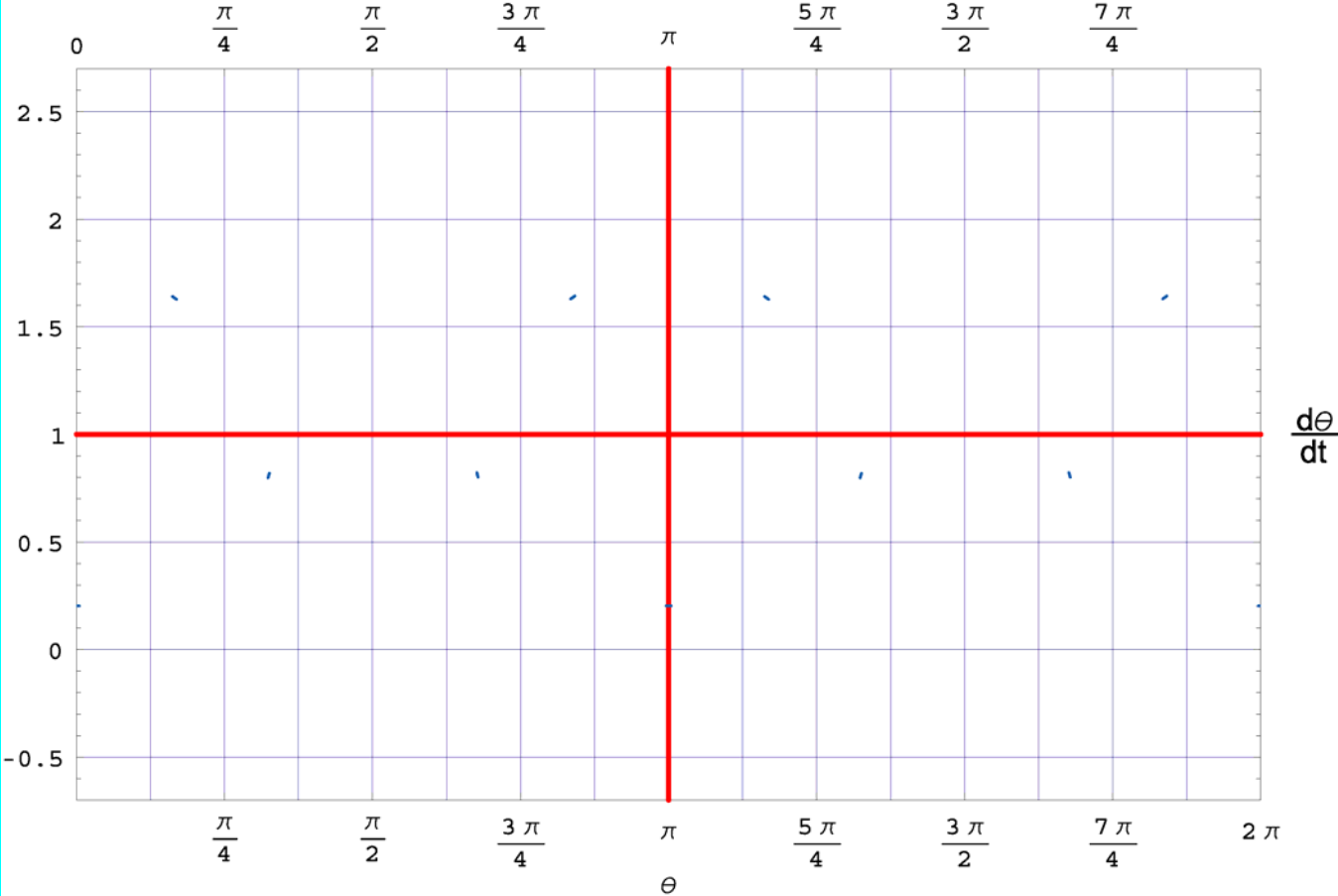




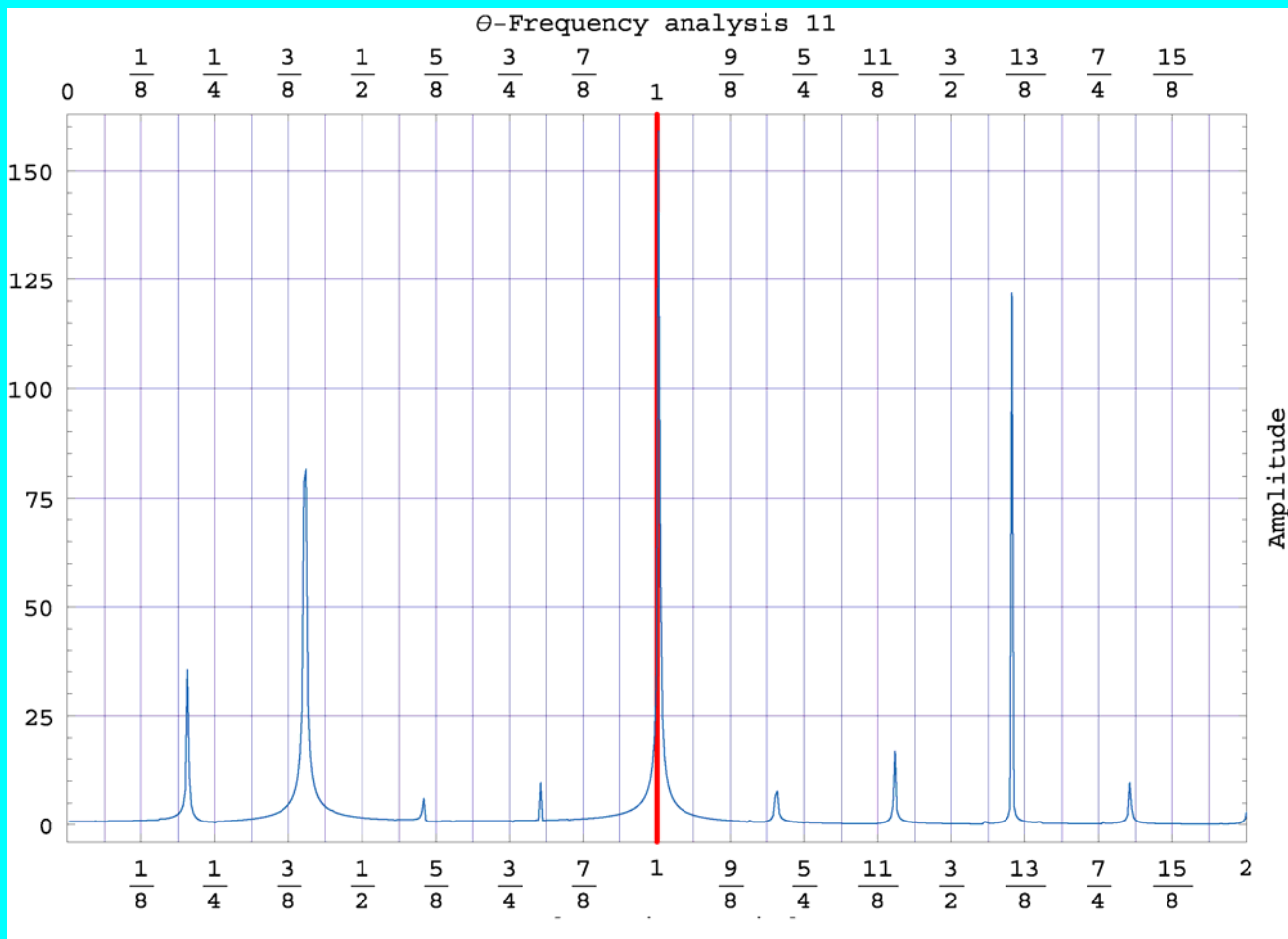
Phobos - Limb profile 5



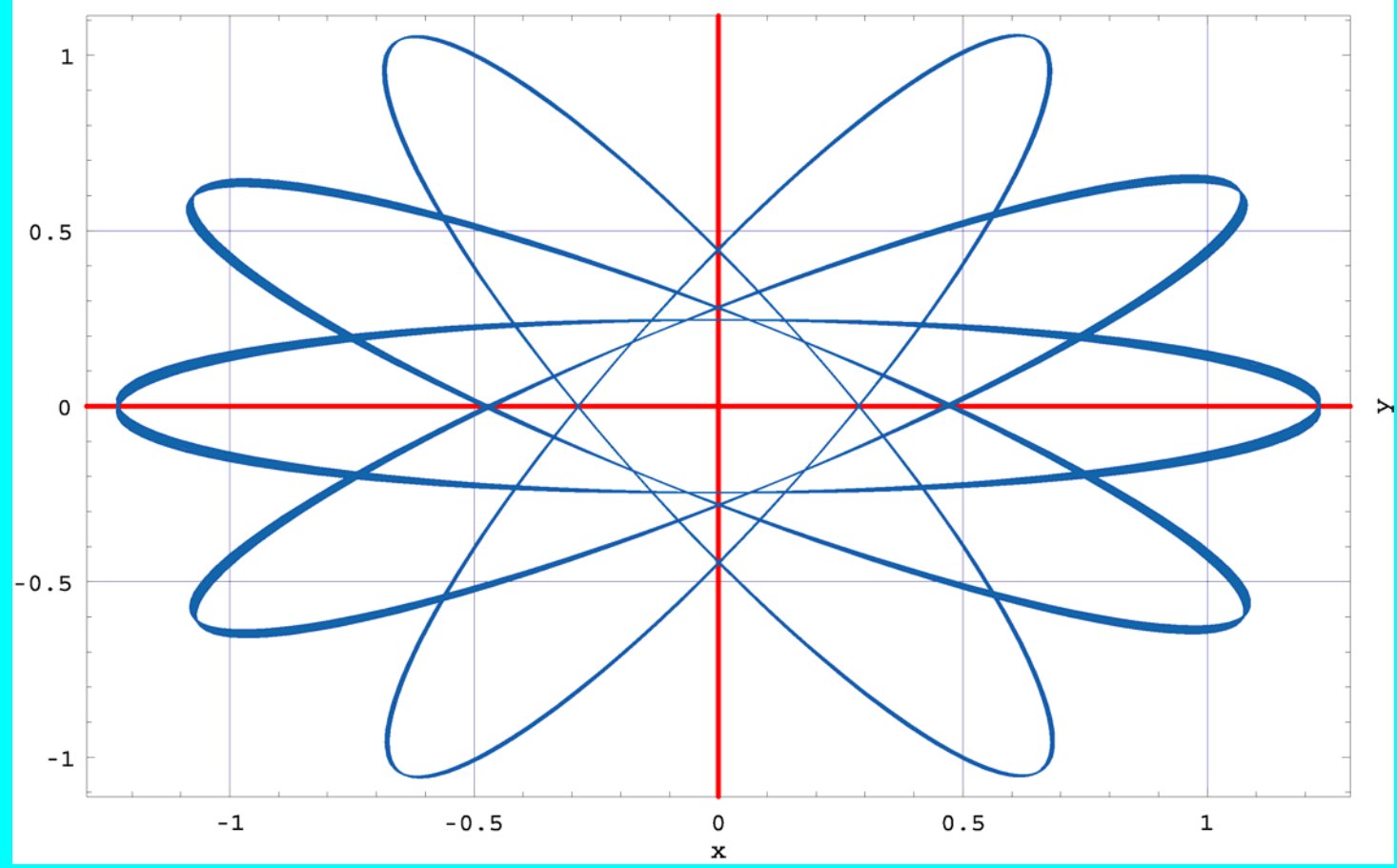
Phase Space 11, Phobos



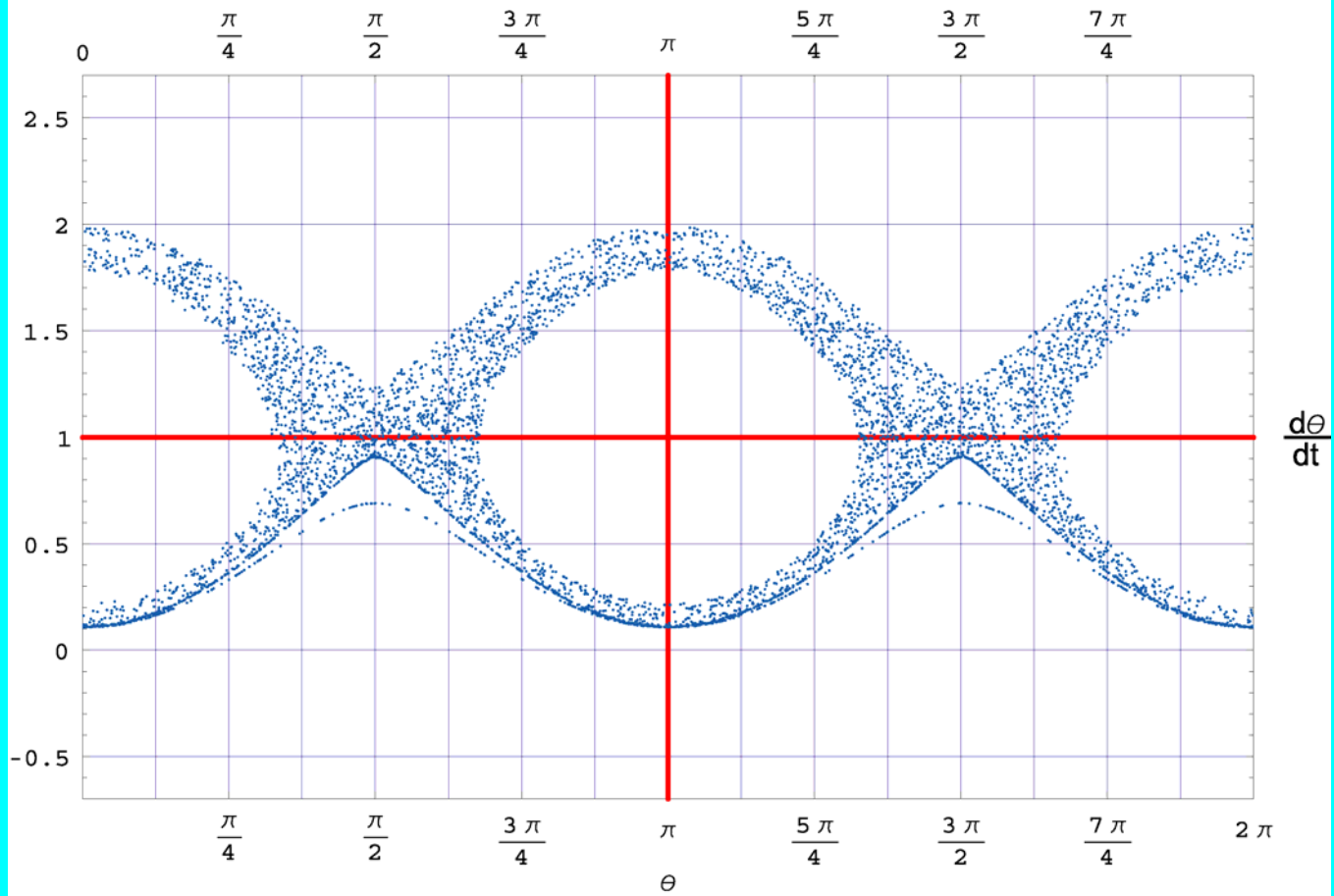


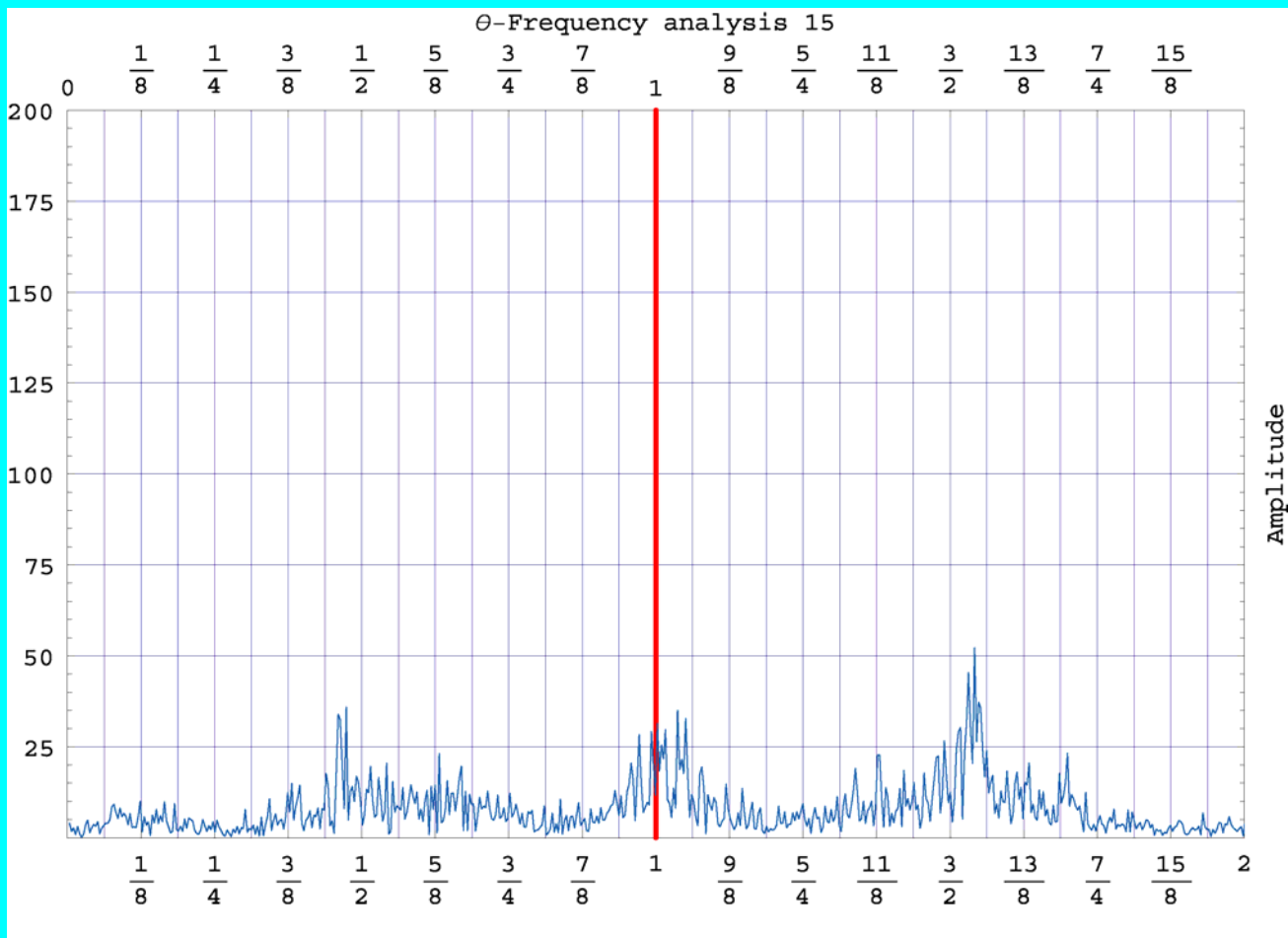


Phobos - Limb profile 11

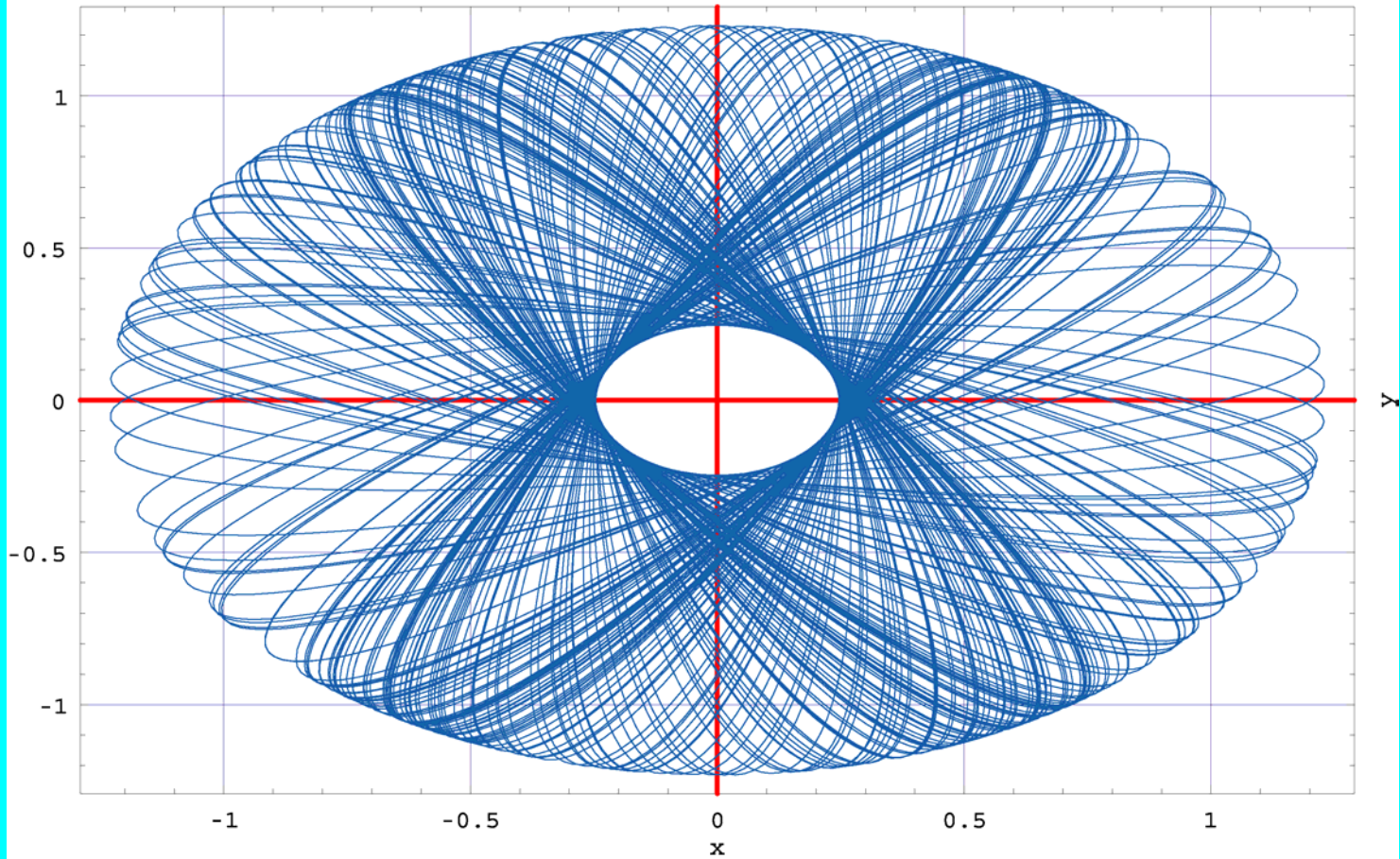


Phase Space 15, Phobos





Phobos - Limb profile 15



# Lyapunov Exponents

$$\frac{d^2\theta}{dt^2} + \frac{3}{2} \frac{B-A}{C} n^2 \left(\frac{a}{r}\right)^3 \sin(2(\theta - f)) = W$$

- Linearized Equations

```

Tmp2 := y3 + C1 * Sin[y3] + C2 * Sin[2 * y3] + C3 * Sin[3 * y3];
Tmp := ((s * (1 - e ^ 2)) / (1 + e * Cos[Tmp2]));

f1 = y2;
f2 = -(1 / 2) * ALPHA ^ 2 * n ^ 2 * s ^ 3 * Tmp ^ -3 * Sin[2 * (y1 - Tmp2)];
f3 = 2 * Pi / T;

MatrixForm[{{Hold[D[f1, y1]], Hold[D[f1, y2]], Hold[D[f1, y3]]},
  {Hold[D[f2, y1]], Hold[D[f2, y2]], Hold[D[f2, y3]]},
  {Hold[D[f3, y1]], Hold[D[f3, y2]], Hold[D[f3, y3]]}} .
  {dx, dy, dz}]

{dx Hold[∂y1 f1] + dy Hold[∂y2 f1] + dz Hold[∂y3 f1]
{dx Hold[∂y1 f2] + dy Hold[∂y2 f2] + dz Hold[∂y3 f2]
{dx Hold[∂y1 f3] + dy Hold[∂y2 f3] + dz Hold[∂y3 f3]

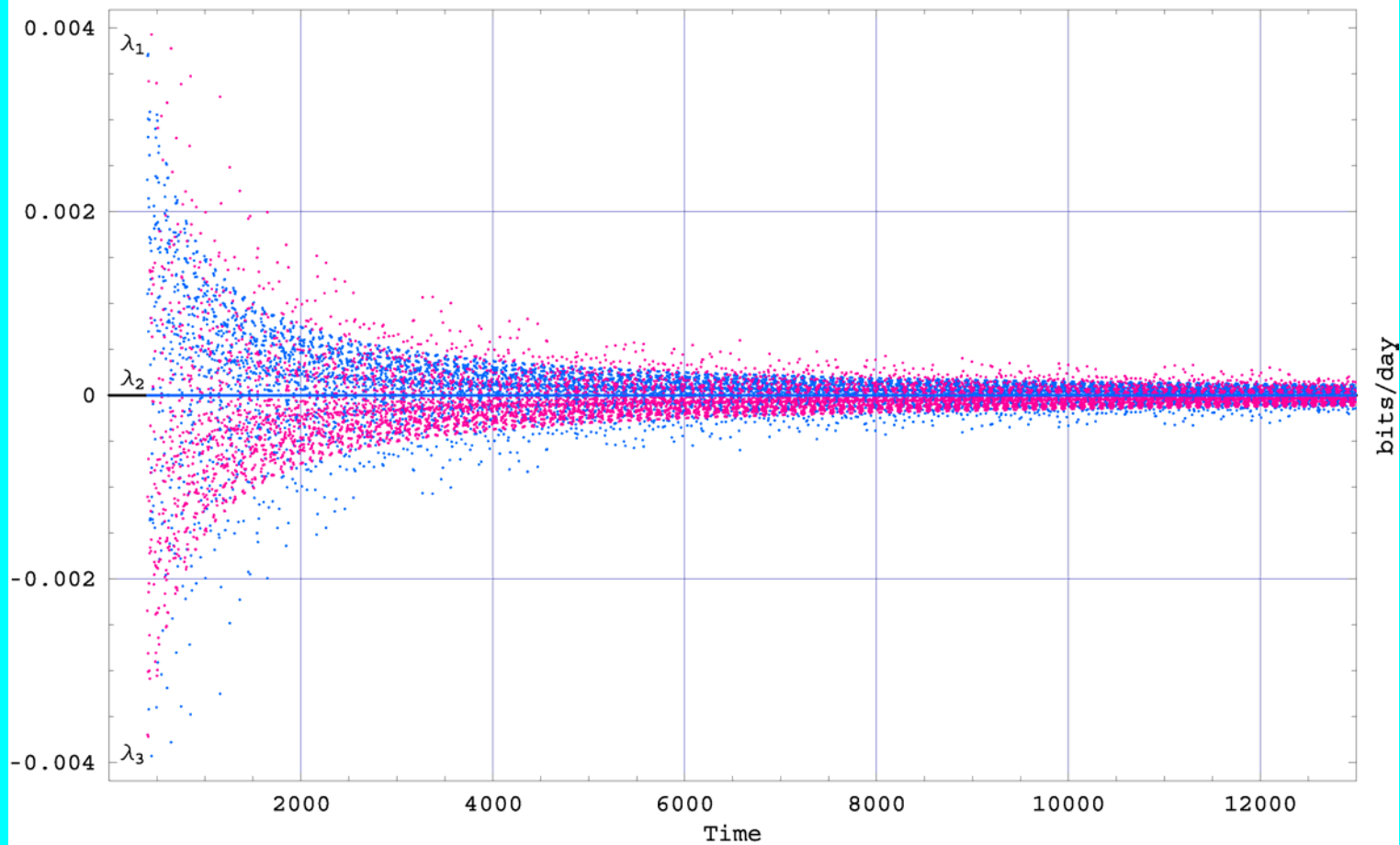
k = ReleaseHold[%];

i = Simplify[k]

{dy,  $\frac{1}{4 (-1 + e^2)^3} (\text{ALPHA}^2 n^2 (1 + e \text{Cos}[y3 + C1 \text{Sin}[y3] + C2 \text{Sin}[2 y3] + C3 \text{Sin}[3 y3]])^2$ 
(-dz (1 + C1 Cos[y3] + 2 C2 Cos[2 y3] + 3 C3 Cos[3 y3])
(5 e Cos[2 y1 - 3 y3 - 3 C1 Sin[y3] - 3 C2 Sin[2 y3] - 3 C3 Sin[3 y3]] +
4 Cos[2 (y1 - y3 - C1 Sin[y3] - C2 Sin[2 y3] - C3 Sin[3 y3]]) -
e Cos[2 y1 - y3 - C1 Sin[y3] - C2 Sin[2 y3] - C3 Sin[3 y3]]) +
4 dx Cos[2 (y1 - y3 - C1 Sin[y3] - C2 Sin[2 y3] - C3 Sin[3 y3]])
(1 + e Cos[y3 + C1 Sin[y3] + C2 Sin[2 y3] + C3 Sin[3 y3]])},
0}

```

Lyapunov-exponents



Lyapunov exponents

